

HANDS-ON TUTORIAL WORKSHOP, JULY 19TH 2011

THERMOSTATS AND THERMAL TRANSPORT IN SOLIDS

Christian Carbogno



MAX-PLANCK-GESELLSCHAFT

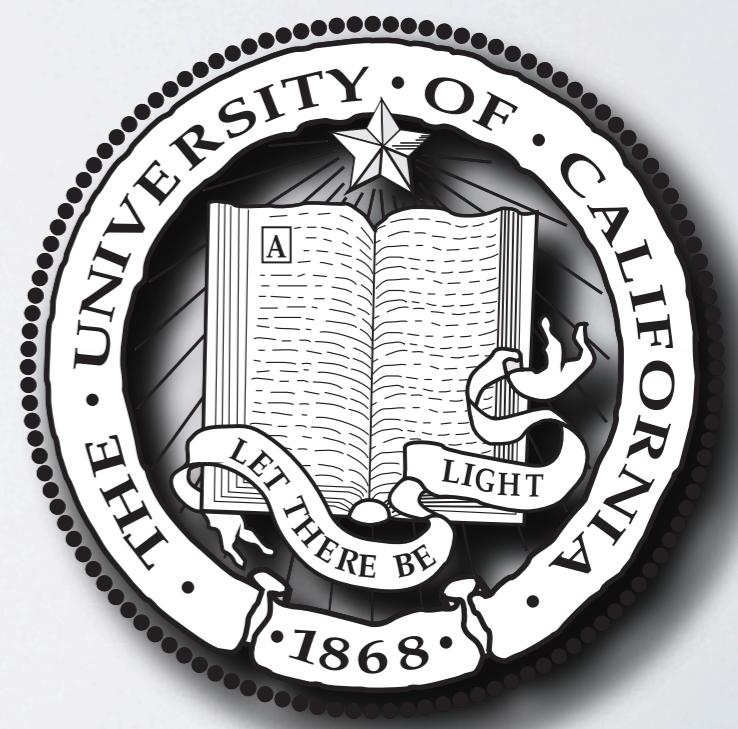
FRITZ-HABER-INSTITUT
DER MAX-PLANCK-GESELLSCHAFT,
BERLIN - GERMANY

HANDS-ON TUTORIAL WORKSHOP, JULY 19TH 2011

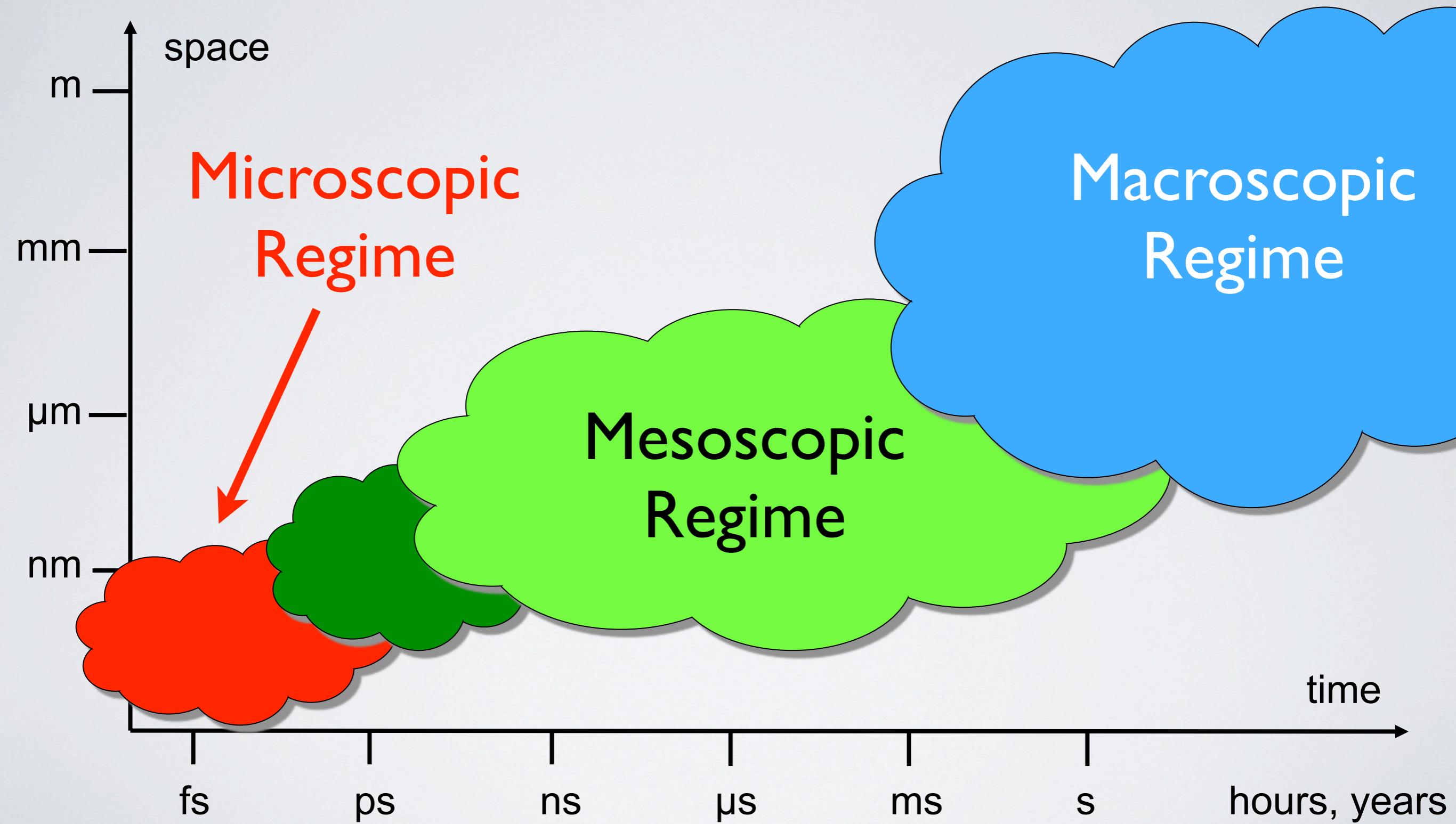
THERMOSTATS AND THERMAL TRANSPORT IN SOLIDS

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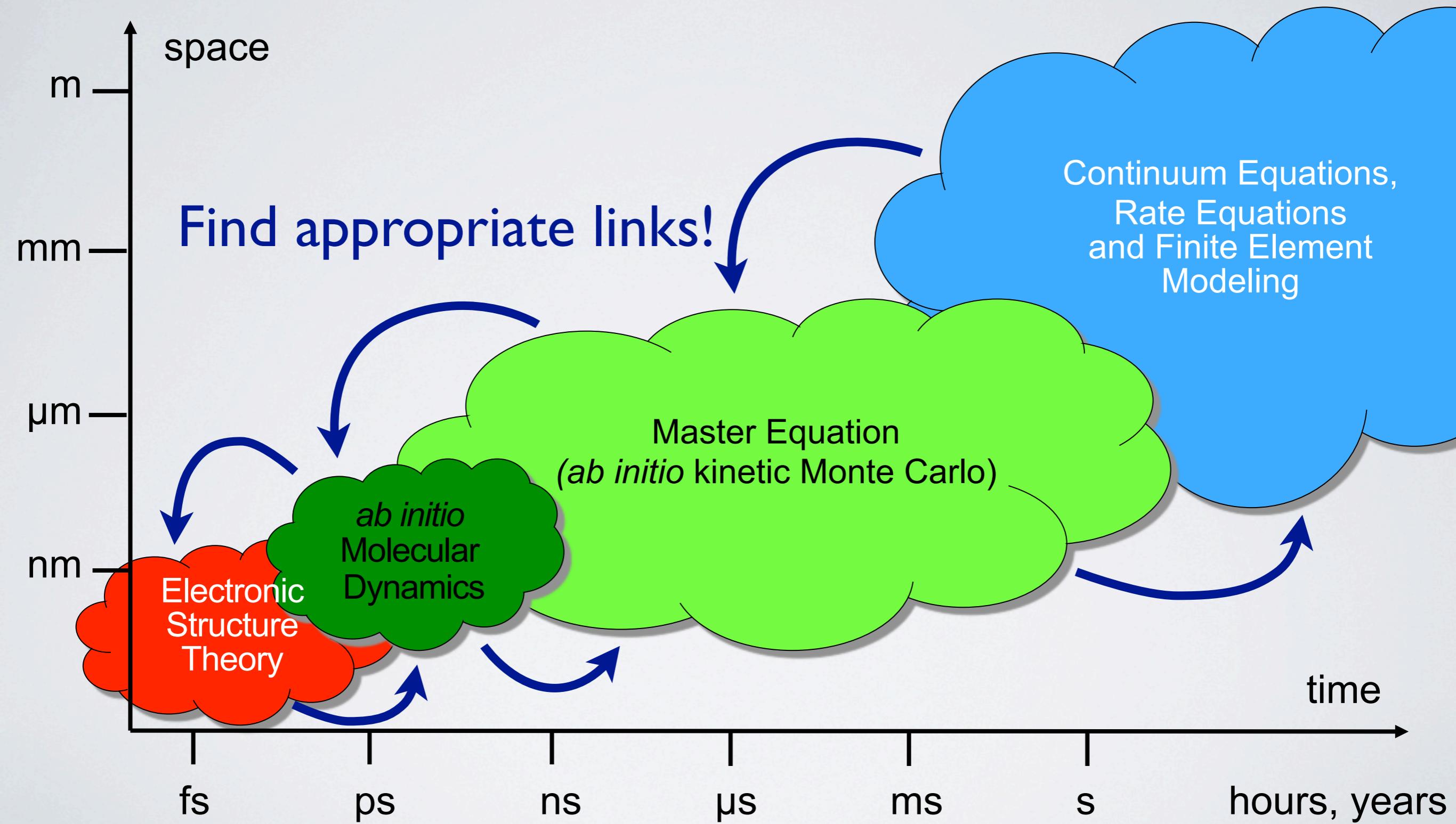
MATERIALS DEPARTMENT,
UNIVERSITY OF CALIFORNIA
SANTA BARBARA



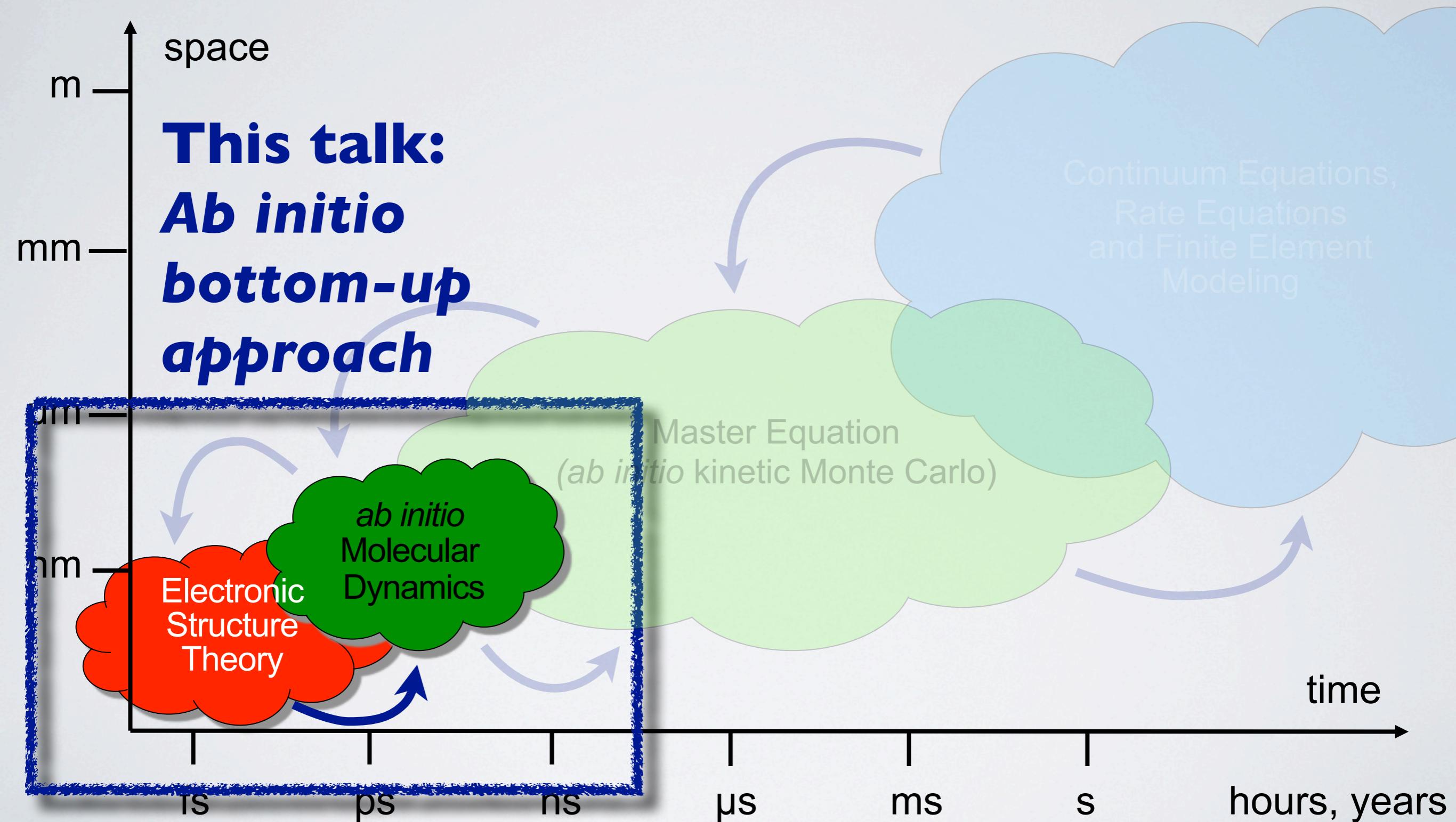
MULTI-SCALE MODELING



MULTI-SCALE MODELING



MULTI-SCALE MODELING



THERMODYNAMIC EQUILIBRIUM IN A NUTSHELL

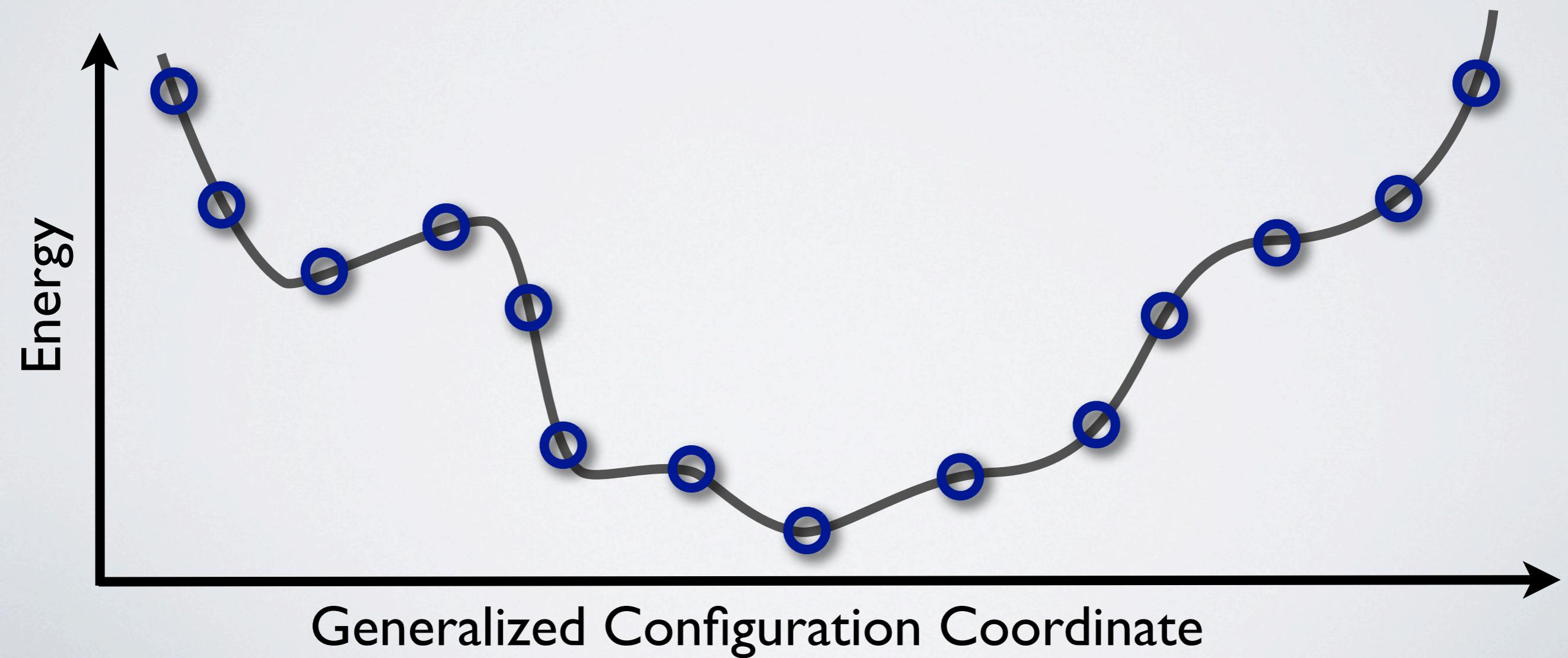
A classical system with **N** atoms, a **Volume V** and a **Temperature T** is described by its **canonical partition function Z(T,V,N)** viz. its **Helmholtz Free Energy F(T,V,N)**.

$$\begin{aligned} F(T, V, N) &= -k_B \ln (Z(T, V, N)) \\ &= -k_B \ln \left(\frac{1}{N! \hbar^{3N}} \int \exp \left(-\frac{\mathcal{H}(\{\mathbf{x}_i\}, \{\mathbf{p}_i\})}{k_B T} \right) \{d^3 \mathbf{x}_i\} \{d^3 \mathbf{p}_i\} \right) \end{aligned}$$


Calculating the energy of **all possible configurations $\{\mathbf{x}_i\}$** is **numerically unfeasible** even for very small systems.

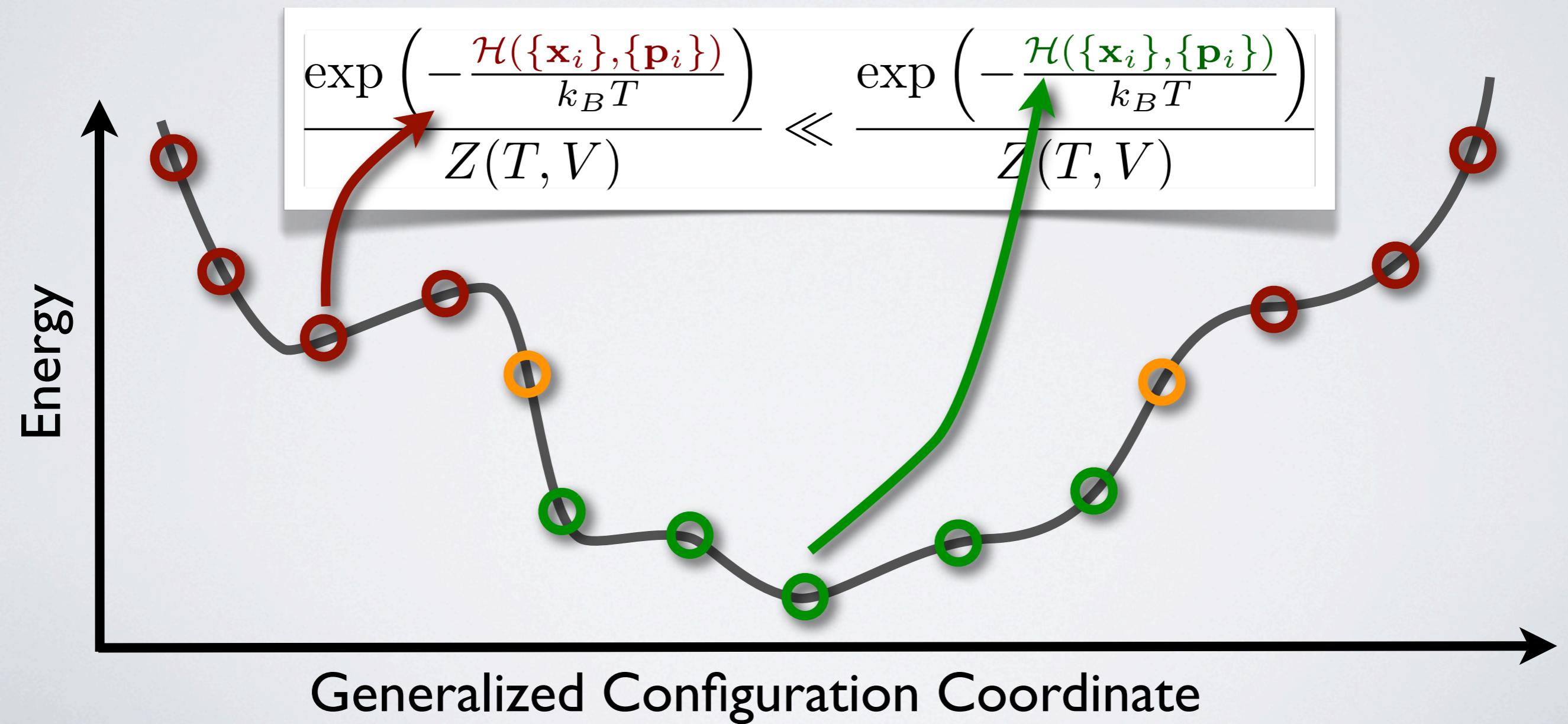
THERMODYNAMIC AVERAGE

$$\langle X \rangle_{T,V} = \frac{1}{Z(T,V)} \int X \exp\left(-\frac{\mathcal{H}(\{\mathbf{x}_i\}, \{\mathbf{p}_i\})}{k_B T}\right) \{d^3\mathbf{x}_i\} \{d^3\mathbf{p}_i\}$$



THERMODYNAMIC AVERAGE

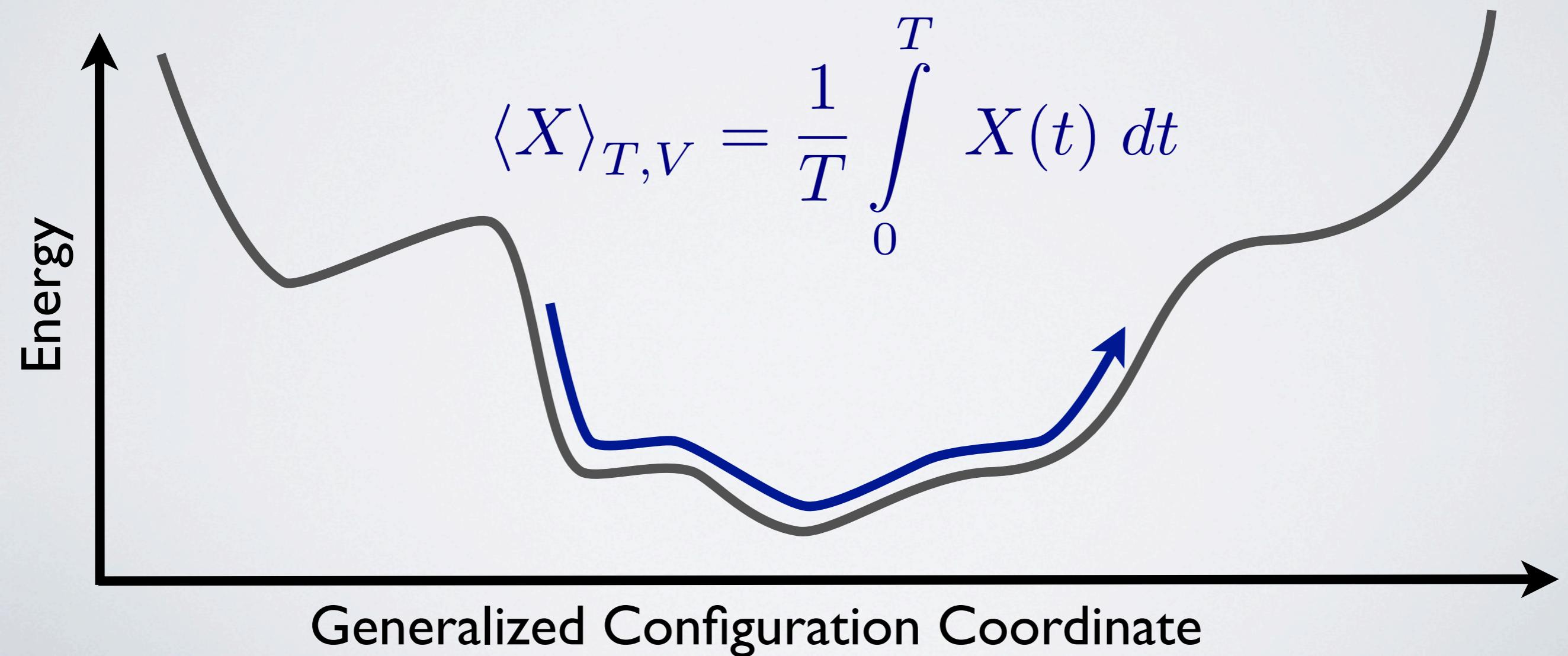
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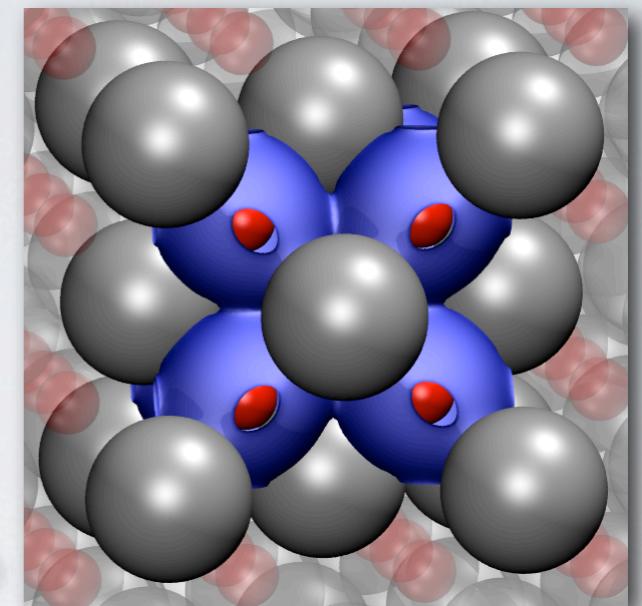
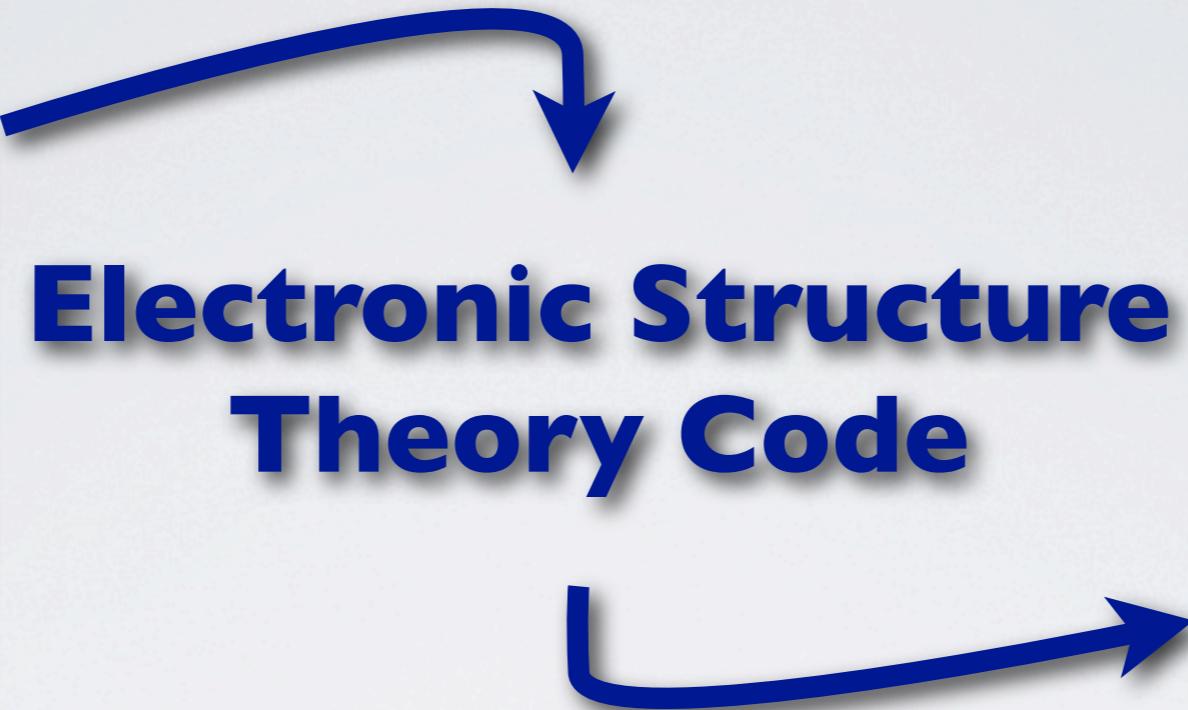
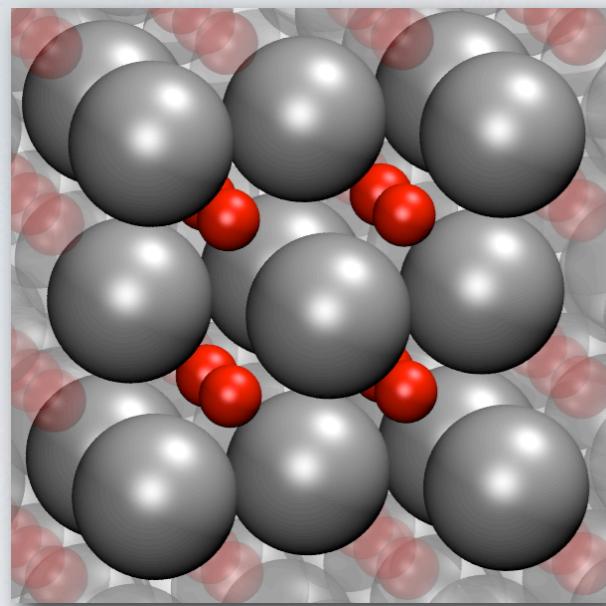
ERGODIC HYPOTHESIS

All accessible micro-states are **equiprobable**
over a long period of time:

The Time Average is equal to the Ensemble Average!



AB INITIO MOLECULAR DYNAMICS



Input:
Geometry, Species

Output:
total energy & forces

Update geometry

Iterative Approach: Explore the Dynamics of the Atoms!

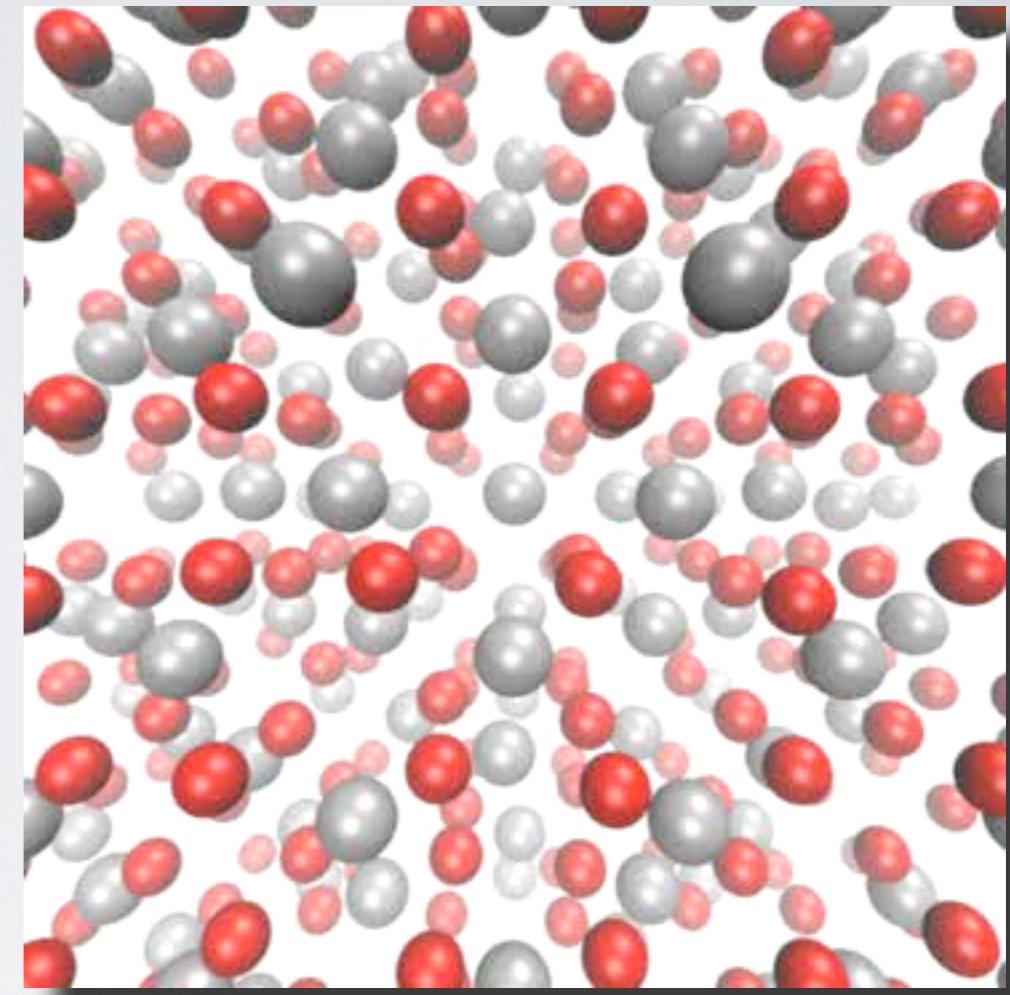
AB INITIO MOLECULAR DYNAMICS

Numerical Integration
of the **equations of motion**

L.Verlet, Phys. Rev. **159**, 98 (1967).

$$M_I \ddot{\mathbf{R}}_I(t) = \mathbf{F}_I(\mathbf{R}_1(t), \dots, \mathbf{R}_N(t))$$

Initial conditions have to
be specified!



The **Verlet Algorithm** conserves the number of particles **N**, the volume **V**, and the energy **E**.

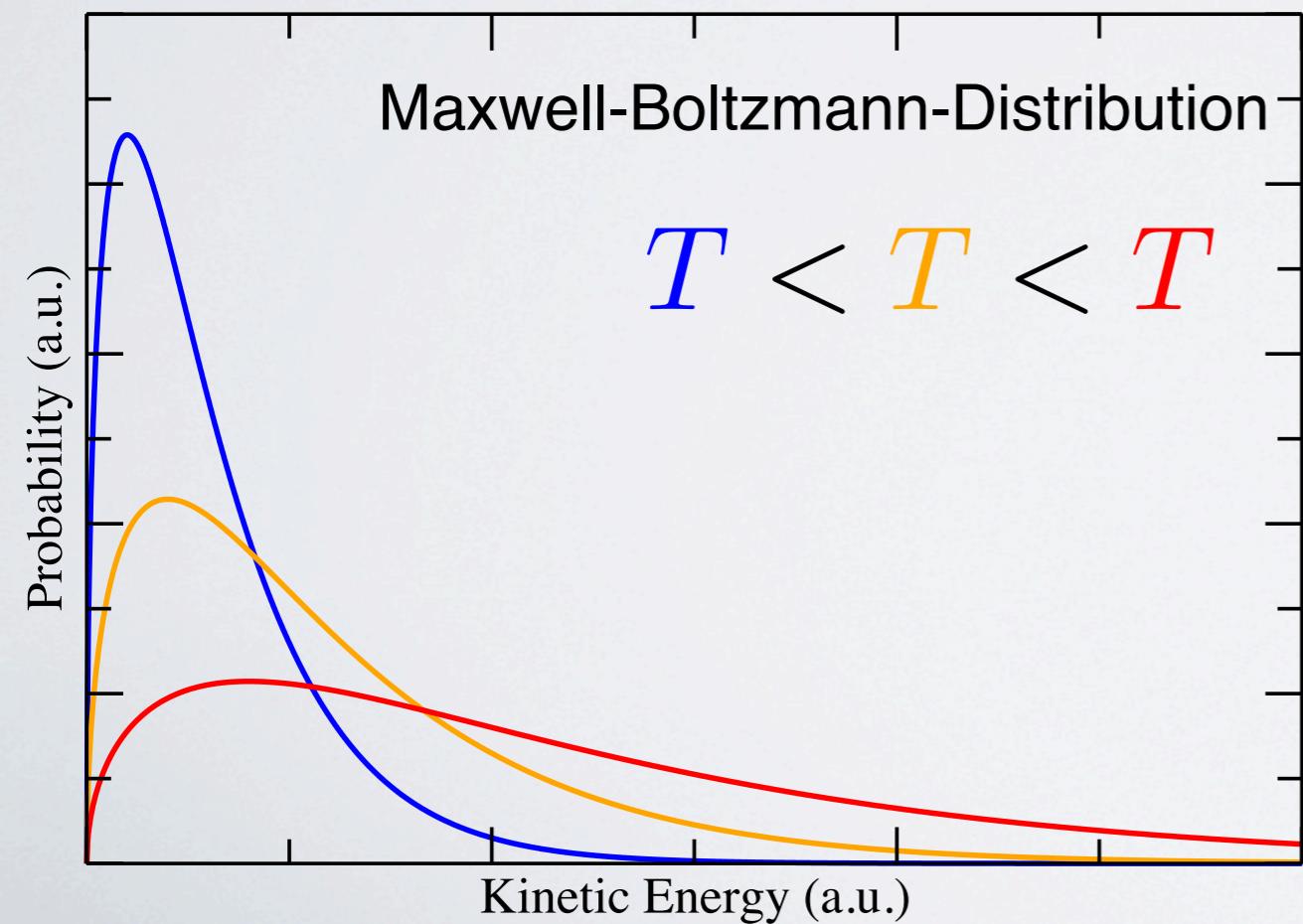
⇒ **Micro-canonical Ensemble**

THE CANONICAL ENSEMBLE



Conserved quantities:

- Number of particles **N**
- Volume **V**
- Average Temperature **T**



$$T = \langle T(t) \rangle = \left\langle \frac{2 E_{kin}}{3 k_B N_{dof}} \right\rangle$$
$$\sigma_T^2 = \frac{2 \langle T(t) \rangle^2}{3 N_{dof}}$$

kinetic Energy
↓
deg. of freedom ↑

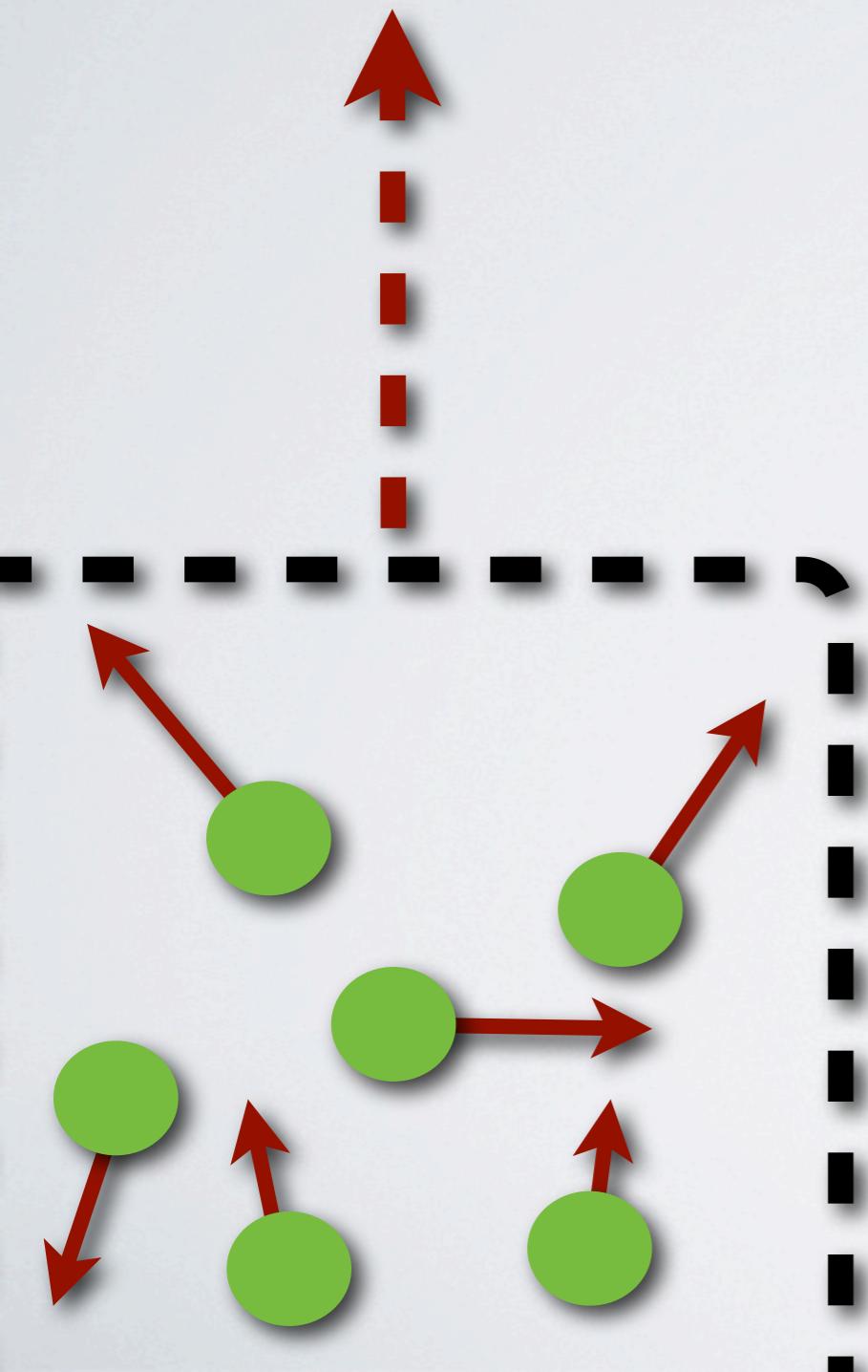
The equation for average temperature T shows it as the expectation value of the kinetic energy per degree of freedom. The standard deviation σ_T^2 is shown as the variance of the temperature, which is proportional to the square of the average temperature divided by the number of degrees of freedom.

I. Thermostats for the Canonical Ensemble: A Historical Overview

BERENDSEN THERMOSTAT

H. J. C. Berendsen, et al. *J. Chem. Phys.* **81** 3684 (1984).

I. Calculate **instantaneous temperature $T(t_0)$** of the system



BERENDSEN THERMOSTAT

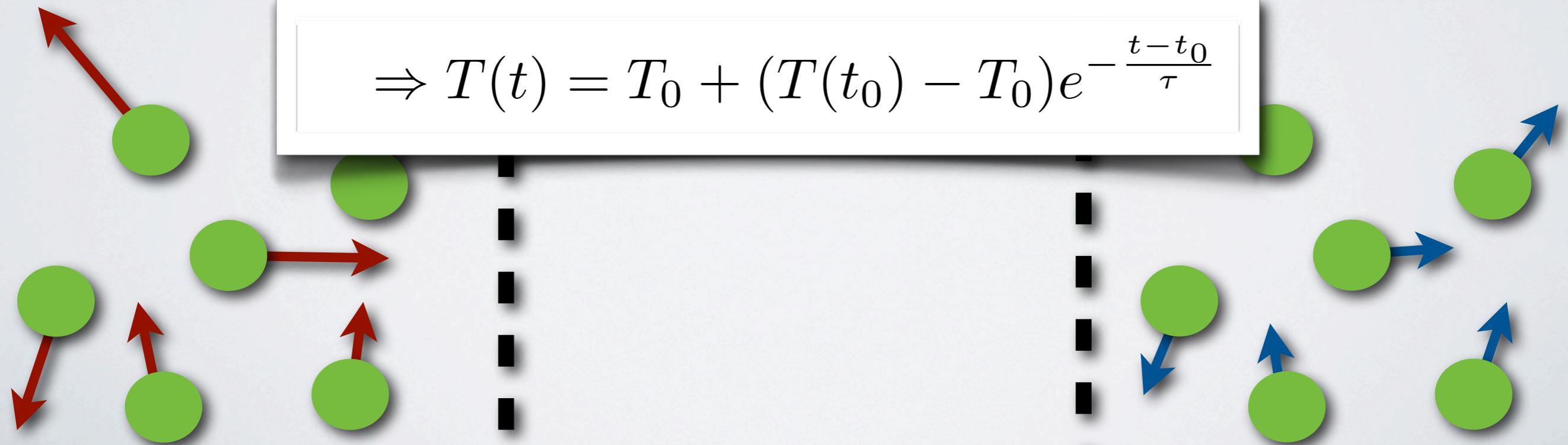
H. J. C. Berendsen, et al. *J. Chem. Phys.* **81** 3684 (1984).

I. Calculate **instantaneous temperature $T(t_0)$** of the system

2. **Rescale** the velocities by

$$\lambda = \left[1 + \frac{\Delta t}{\tau} \left(\frac{T_0}{T} - 1 \right) \right]^{1/2}$$

$$\Rightarrow T(t) = T_0 + (T(t_0) - T_0)e^{-\frac{t-t_0}{\tau}}$$



BERENDSEN THERMOSTAT

H. J. C. Berendsen, et al. *J. Chem. Phys.* **81** 3684 (1984).

Pros & Cons:

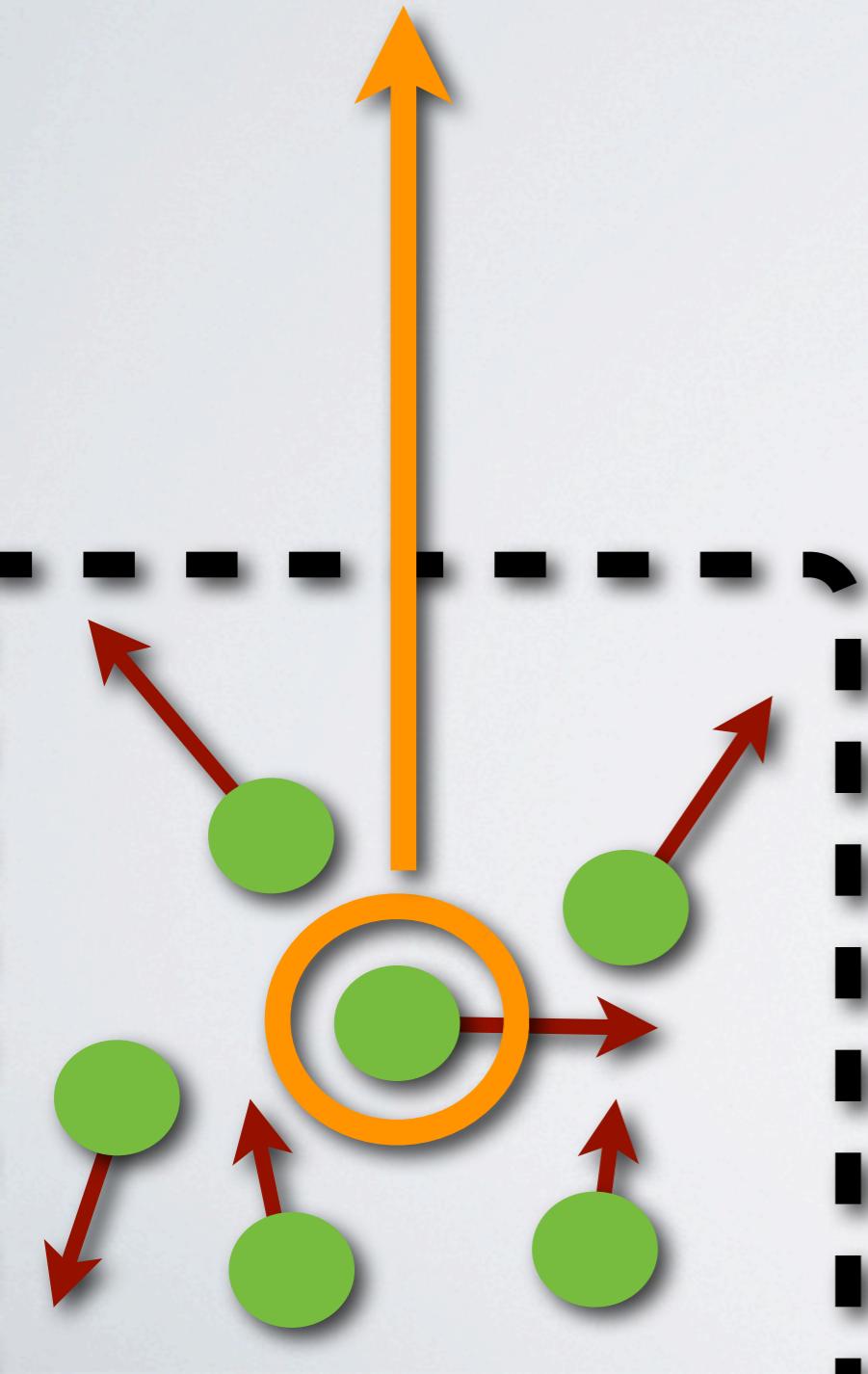
- **Warning:** Suppresses thermal fluctuations
- **Warning:** Does not sample the canonical ensemble
- **Uncertainty:** Sensitive on parameter T

⇒ **Historically important!**

ANDERSEN THERMOSTAT

H. C. Andersen, *J. Chem. Phys.* **72**, 2384 (1980).

- I. Choose a **random atom each v^{th} time step**

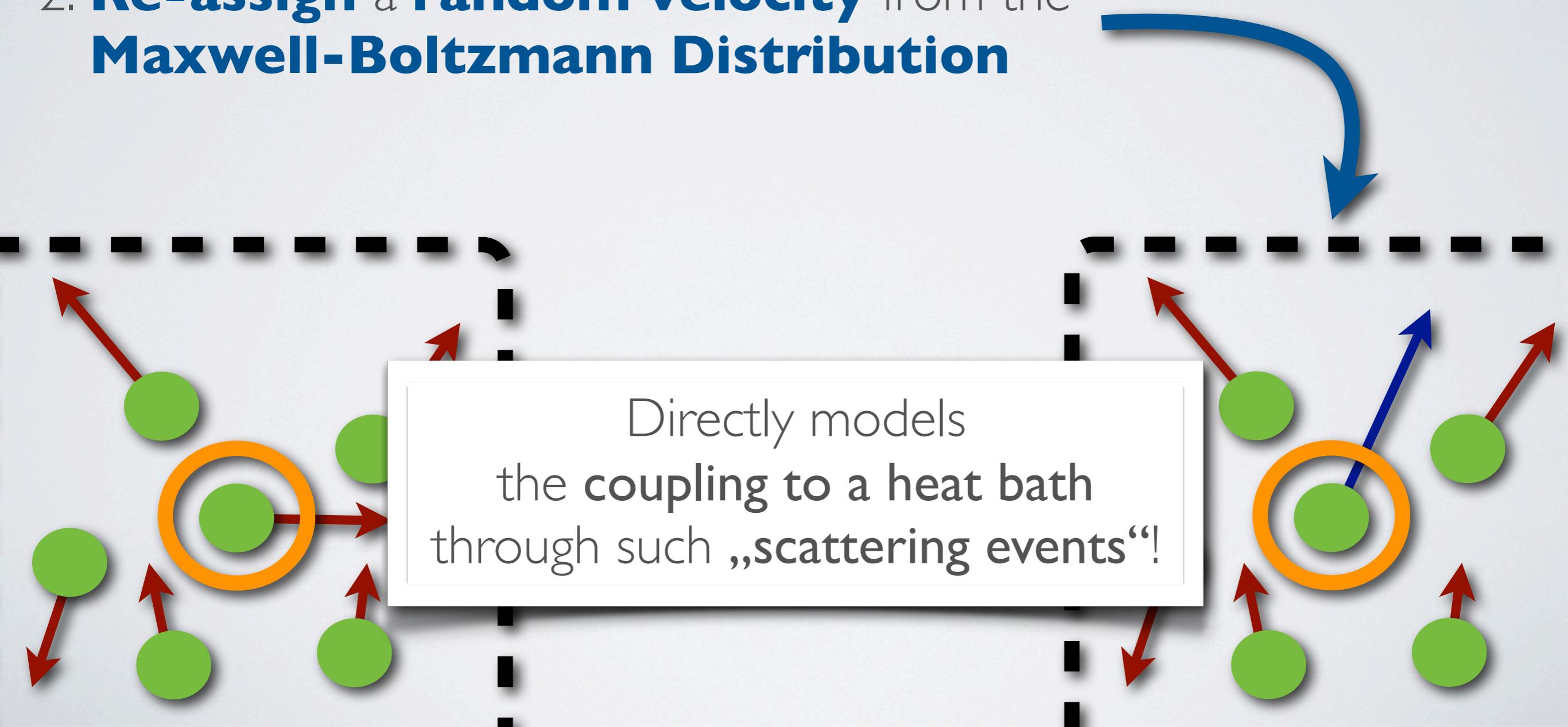


ANDERSEN THERMOSTAT

H. C. Andersen, *J. Chem. Phys.* **72**, 2384 (1980).

- I. Choose a **random atom each v^{th} time step**

2. **Re-assign a random velocity from the Maxwell-Boltzmann Distribution**



ANDERSEN THERMOSTAT

H. C. Andersen, *J. Chem. Phys.* **72**, 2384 (1980).

Pros & Cons:

- Accurate:** Canonical ensemble in principle sampled
- Uncertainty:** Sensitive on parameter \mathbf{v}
- Warning:** Trajectories (Velocities) become discontinuous.

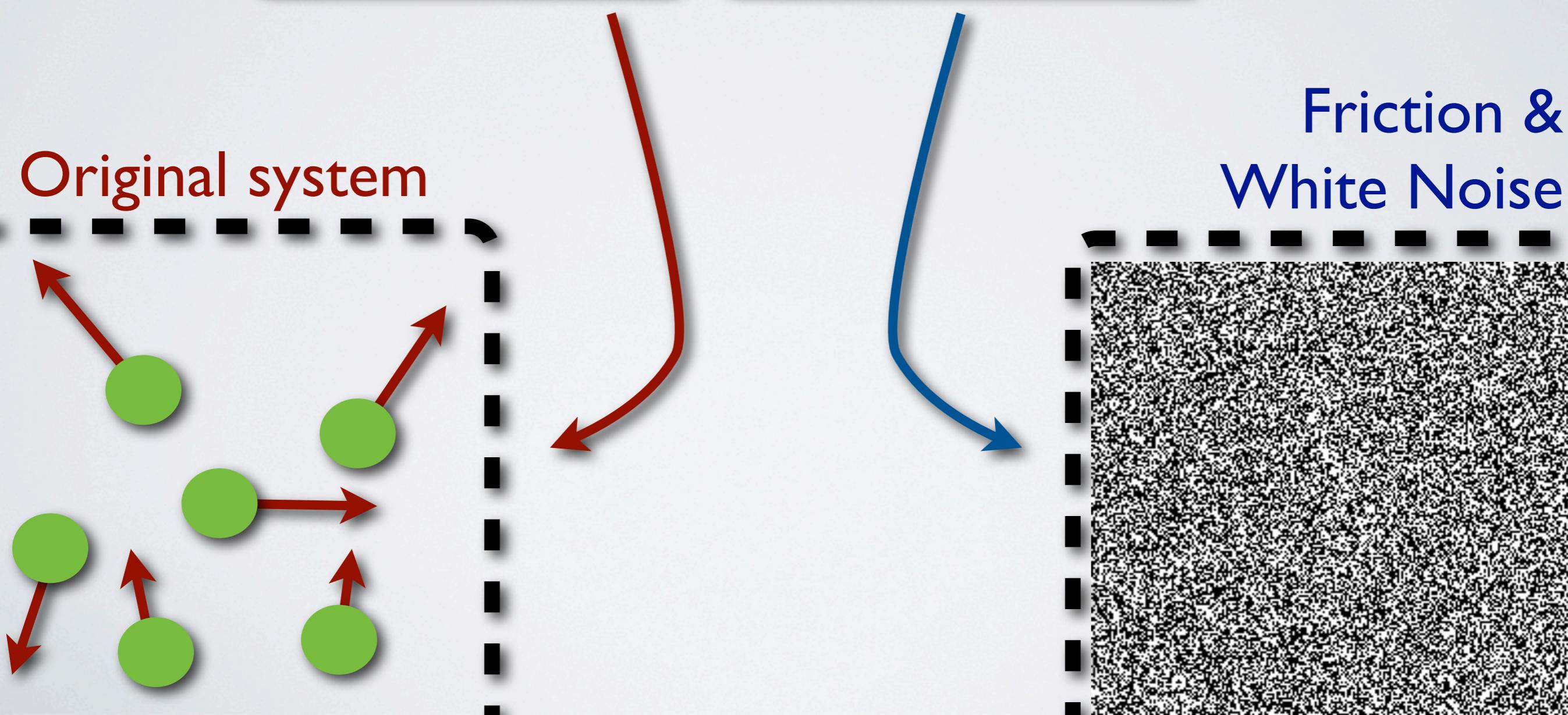
⇒ Historically important!

LANGEVIN'S STOCHASTIC DYNAMICS

S. A. Adelman and J. D. Doll, *J. Chem. Phys.* **64**, 2375 (1976).

Augmented Equations of Motion:

$$m_i \ddot{\mathbf{R}}_i = \mathbf{F}_i - \gamma_i \dot{\mathbf{R}}_i + \Phi_i(t)$$



LANGEVIN'S STOCHASTIC DYNAMICS

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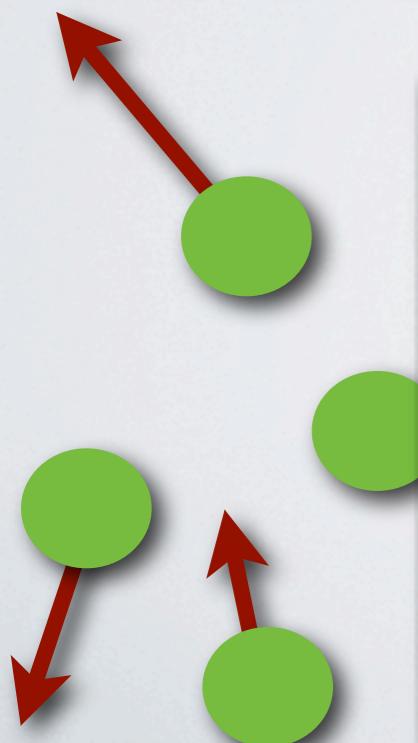
$$m_i \ddot{\mathbf{R}}_i = \mathbf{F}_i - \gamma_i \dot{\mathbf{R}}_i + \Phi_i(t)$$

Original system

Friction &
White Noise

Friction **cools**, Noise **heats**
the system:

$$\langle \Phi_i(t) \cdot \Phi_i(t') \rangle = 6k_B T \gamma_i \delta(t - t')$$



LANGEVIN'S STOCHASTIC DYNAMICS

S. A. Adelman and J. D. Doll, *J. Chem. Phys.* **64**, 2375 (1976).

Pros & Cons:

- Accurate:** Canonical ensemble in principle sampled
- Accurate:** Dynamic properties assessed accurately
- Uncertainty:** Sensitive on parameter(s) γ_i

⇒ Accurate,

but there is no continuity of momentum!

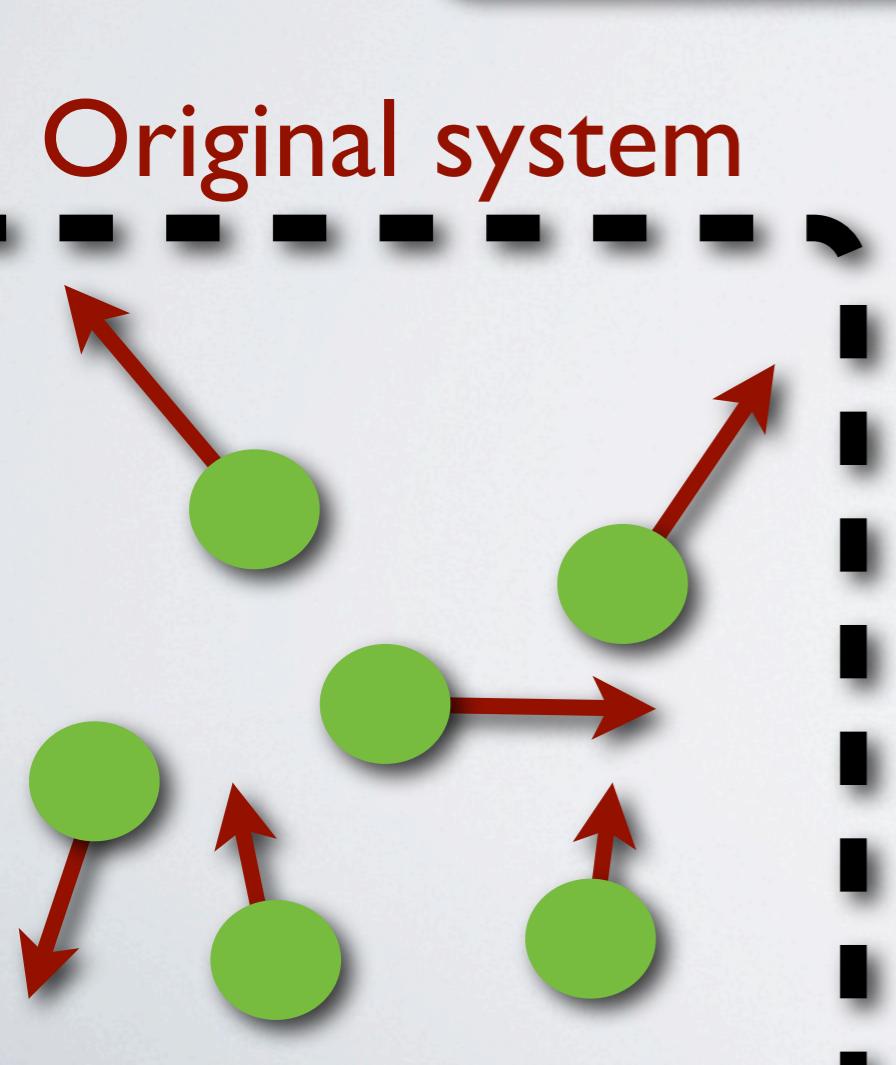
NOSÉ-HOOVER THERMOSTAT

S. Nosé, *J. Chem. Phys.* **81**, 511 (1984) & W. G. Hoover, *Phys. Rev. A* **31**, 1695 (1985).

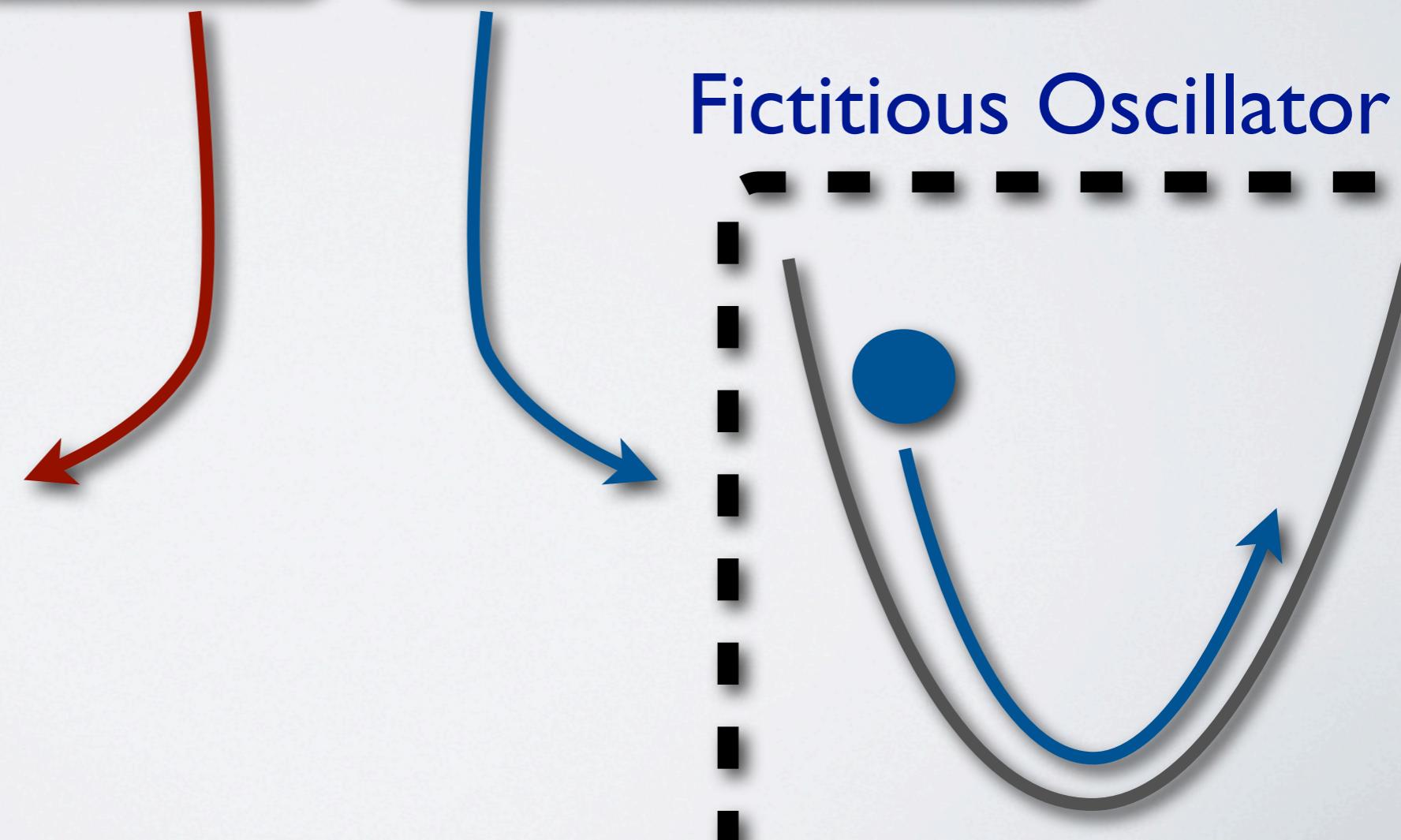
Augmented Hamiltonian:

$$\varepsilon = \sum_i \frac{\mathbf{p}_i^2}{2m_i} + U(\mathbf{r}^N) + \frac{1}{2}Q \dot{\eta} + 3Nk_B T \eta$$

Original system



Fictitious Oscillator



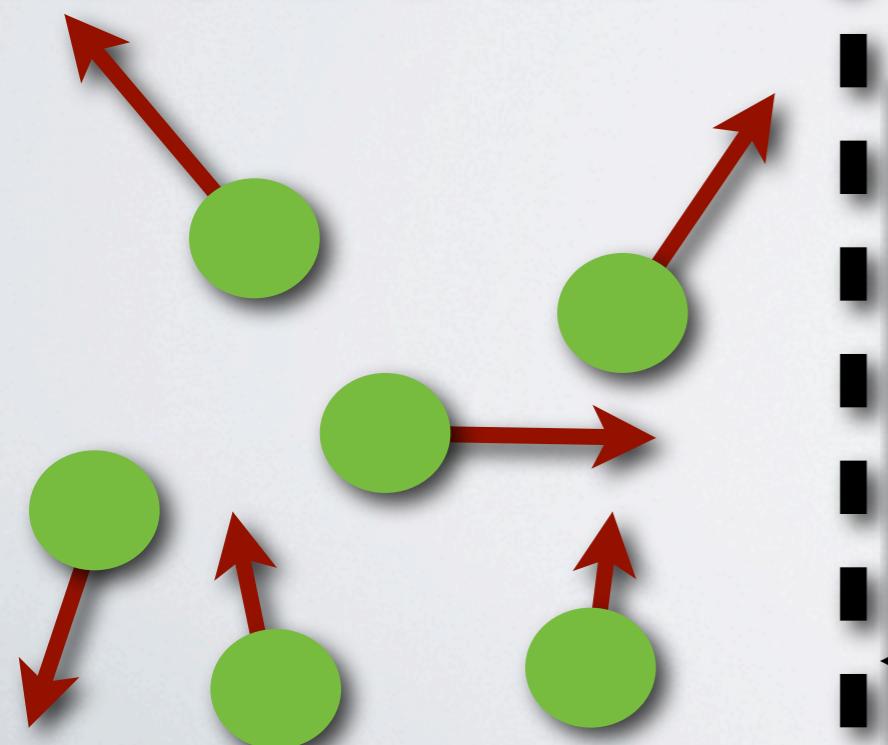
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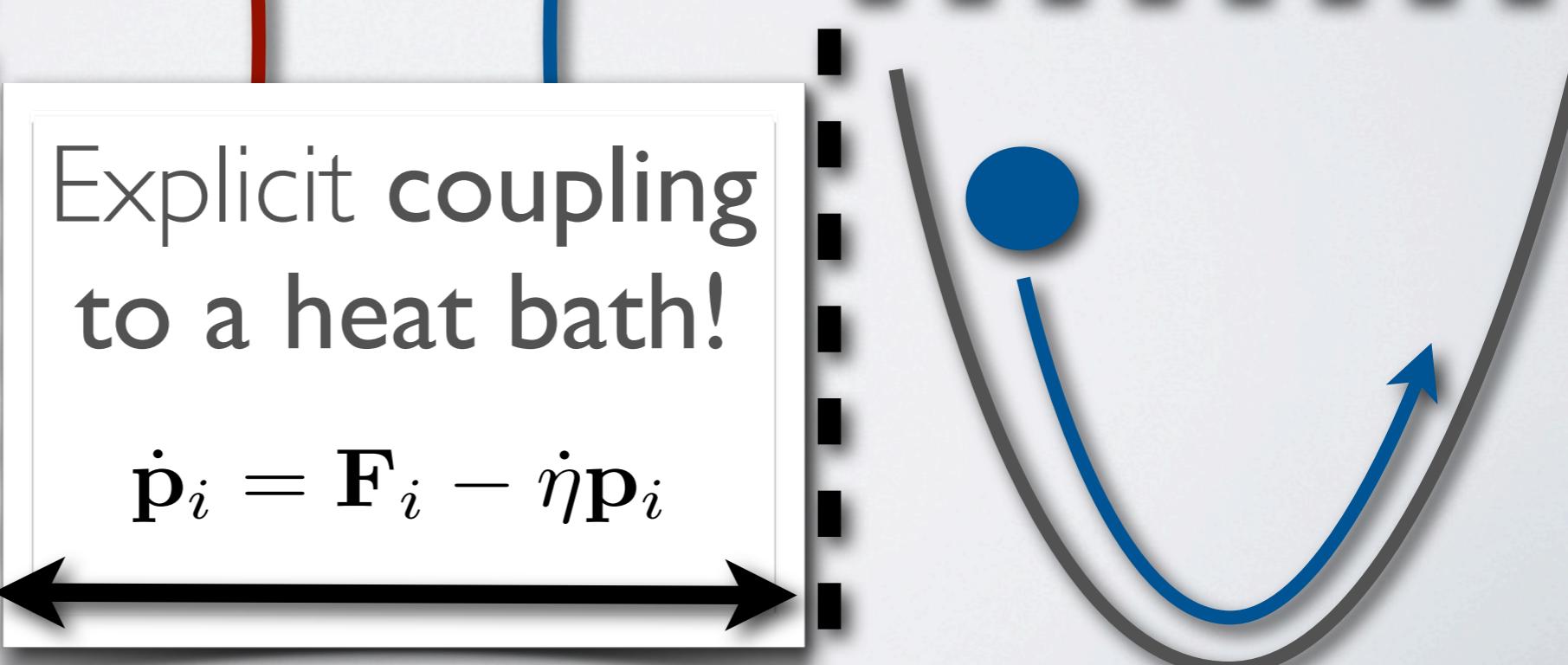
Original system



Fictitious Oscillator

Explicit coupling
to a heat bath!

$$\dot{\mathbf{p}}_i = \mathbf{F}_i - \dot{\eta} \mathbf{p}_i$$



NOSÉ-HOOVER THERMOSTAT

S. Nosé, *J. Chem. Phys.* **81**, 511 (1984) & W. G. Hoover, *Phys. Rev. A* **31**, 1695 (1985).

Pros & Cons:

- Accurate:** Canonical ensemble sampled
- Convenient:** Augmented Total Energy is conserved
- Warning:** Trajectories still „feel“ the harmonic oscillation.
- Sensitivity:** **Q** has to be chosen with care

⇒ Accurate,

but **Q** has to be chosen with care!

BUSSI-DONADIO-PARRINELLO THERMOSTAT

G. Bussi, D. Donadio, and M. Parrinello, *J. Chem. Phys.* **126**, 014101 (2007).

Combine concepts from **velocity rescaling (fast!)** with
concepts from **stochastic thermostats (accurate!).**

Target Temperature follows a Stochastic Diff. Eq.:

$$dT = (\bar{T} - T(t)) \frac{dt}{\tau} + 2\sqrt{\frac{T(t)\bar{T}}{N_f}} \frac{dW(t)}{\sqrt{\tau}}$$

Berendsen'
thermostat

White
Noise

BUSSI-DONADIO-PARRINELLO THERMOSTAT

G. Bussi, D. Donadio, and M. Parrinello, *J. Chem. Phys.* **126**, 014101 (2007).

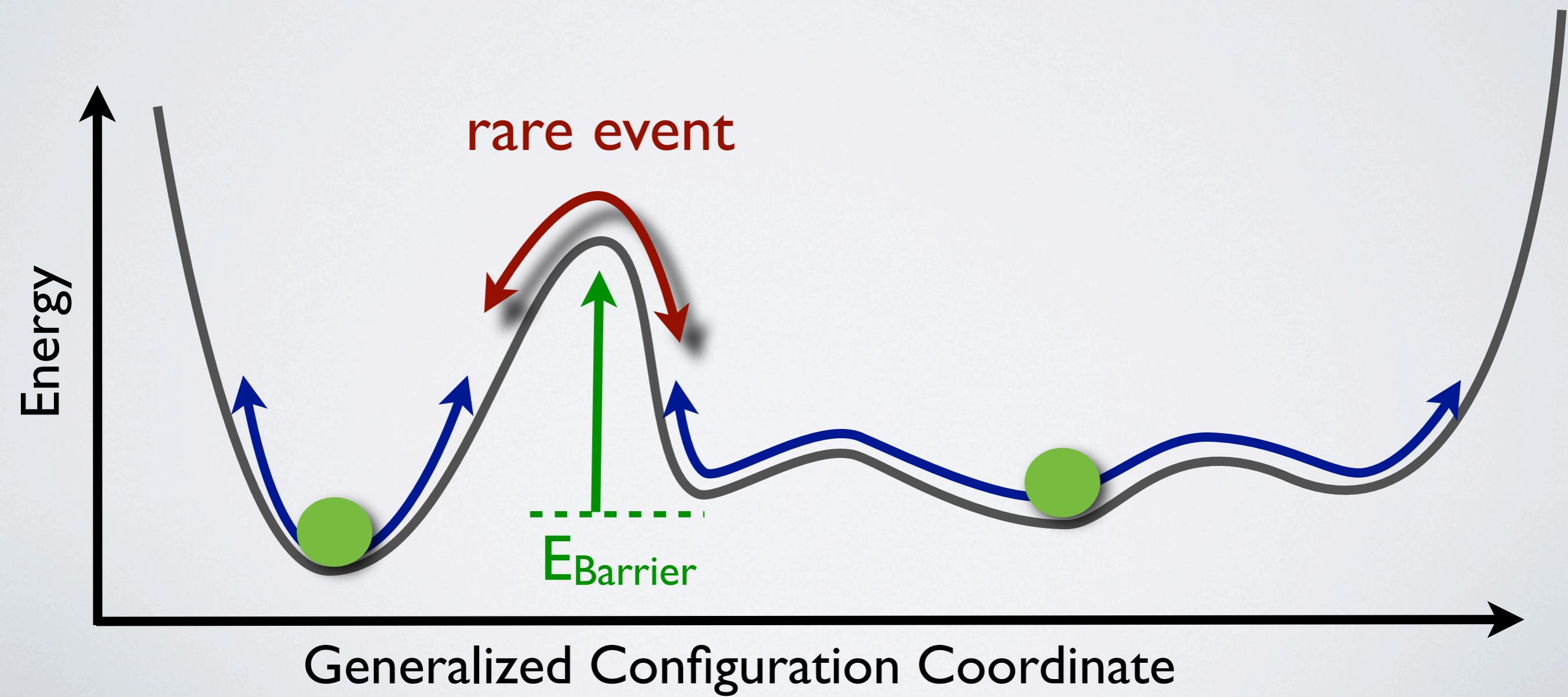
Pros & Cons:

- Accurate:** Canonical ensemble sampled
- Accurate:** Dynamic properties assessed accurately
- Sensitivity:** Almost independent from parameter τ
- Convenient:** Pseudo-Hamiltonian is conserved

⇒ Accurate & very promising,
but also still topic of research!

CAVEAT: RARE EVENTS

Simulating **processes** hindered by **high energetic barriers** can be **extremely expensive** with conventional MD.



CAVEAT: RARE EVENTS

Possible routes:

- Simulated Annealing

e.g., S. Kirkpatrick, C. D. Gelatt, and M. P. Vecchi, *Science* **220**, 671 (1983).

- Accelerated Molecular Dynamics

e.g., A. F. Voter, *Phys. Rev. Lett.* **78**, 3908 (1997).

- Bond Boost

e.g., R. A. Miron, K. A. Fichthorn, *Phys. Rev. Lett.* **93**, 128301 (2004).

- Transition Path Sampling

e.g., P. G. Bolhuis et al. *Annu. Rev. Phys. Chem.* **53**, 291 (2002).

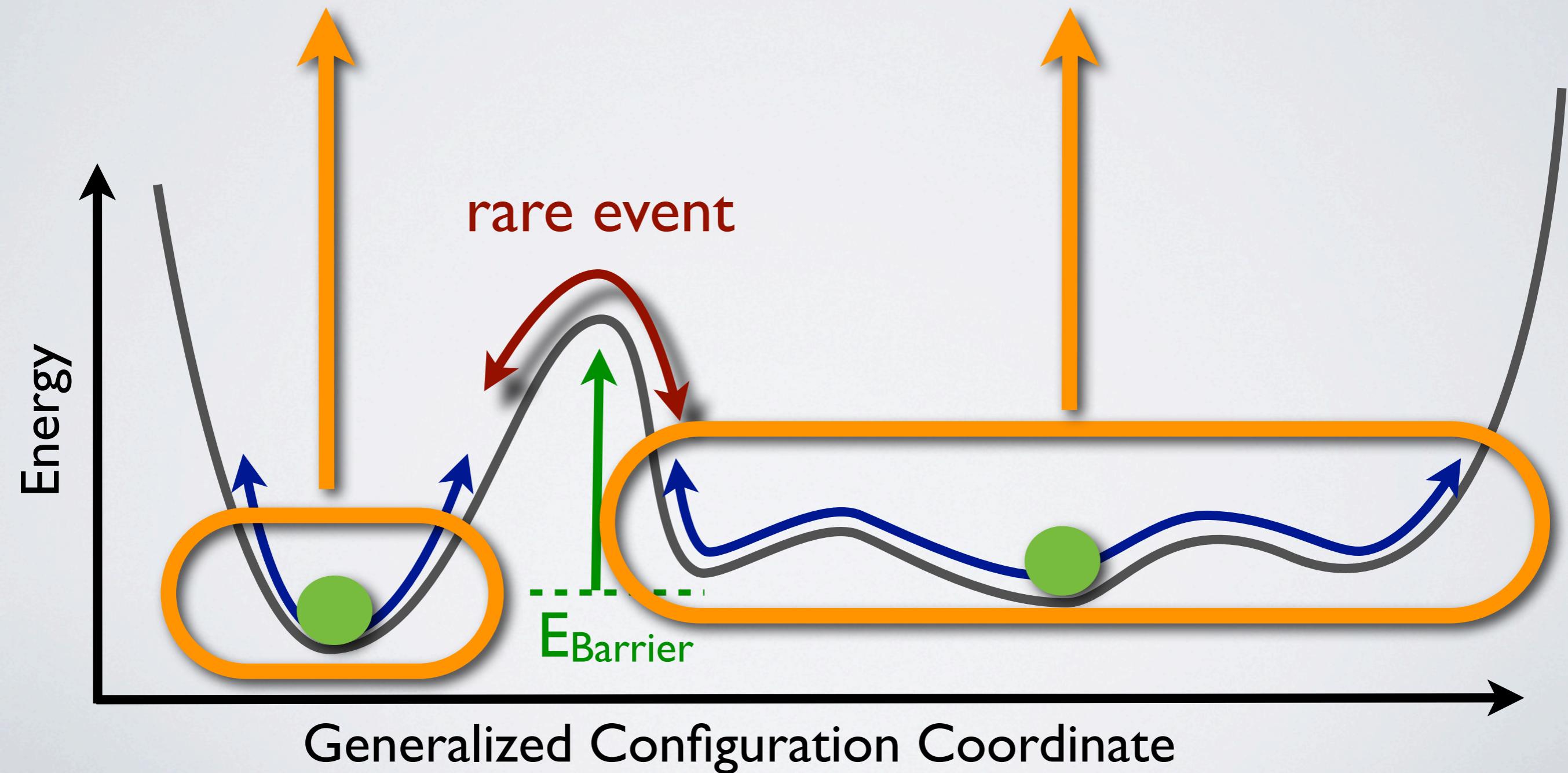
- Kinetic Monte Carlo

e.g., J. Rogal, K. Reuter, and M. Scheffler, *Phys. Rev. B* **77**, 155410 (2008).

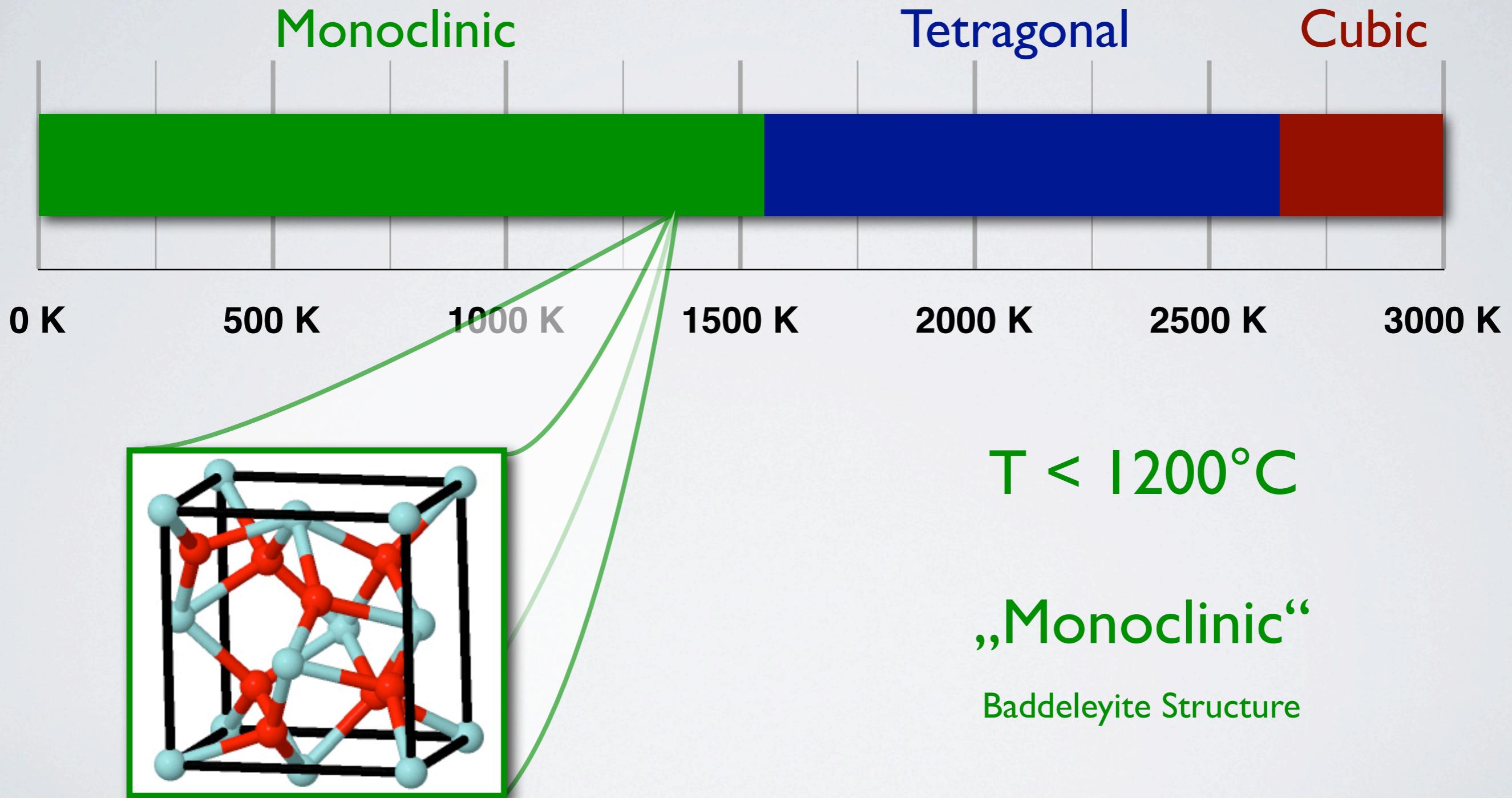
⇒ **Talk on Thursday 21nd**

CAVEAT: RARE EVENTS

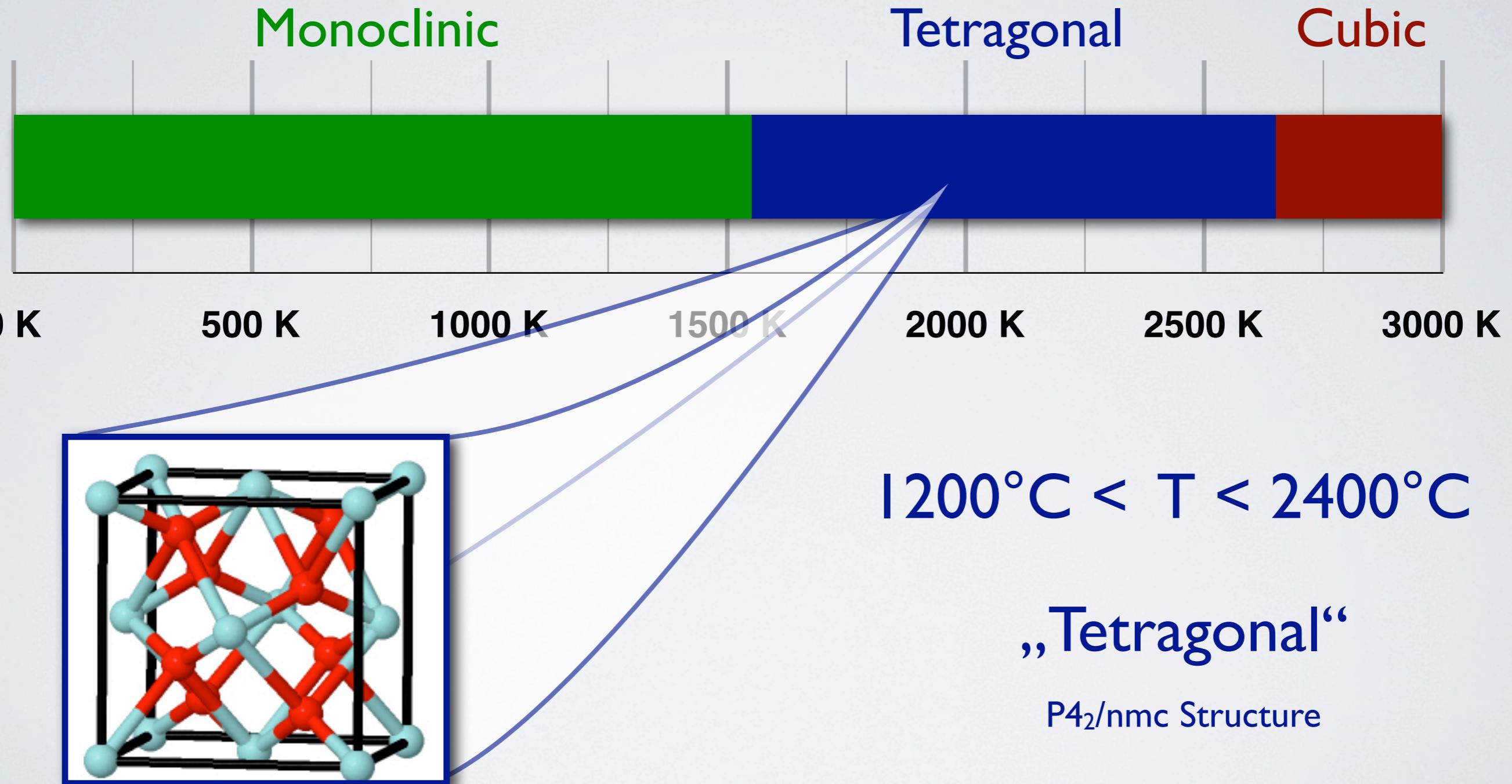
Assess **relative stability**
without simulating the **rare event** itself.



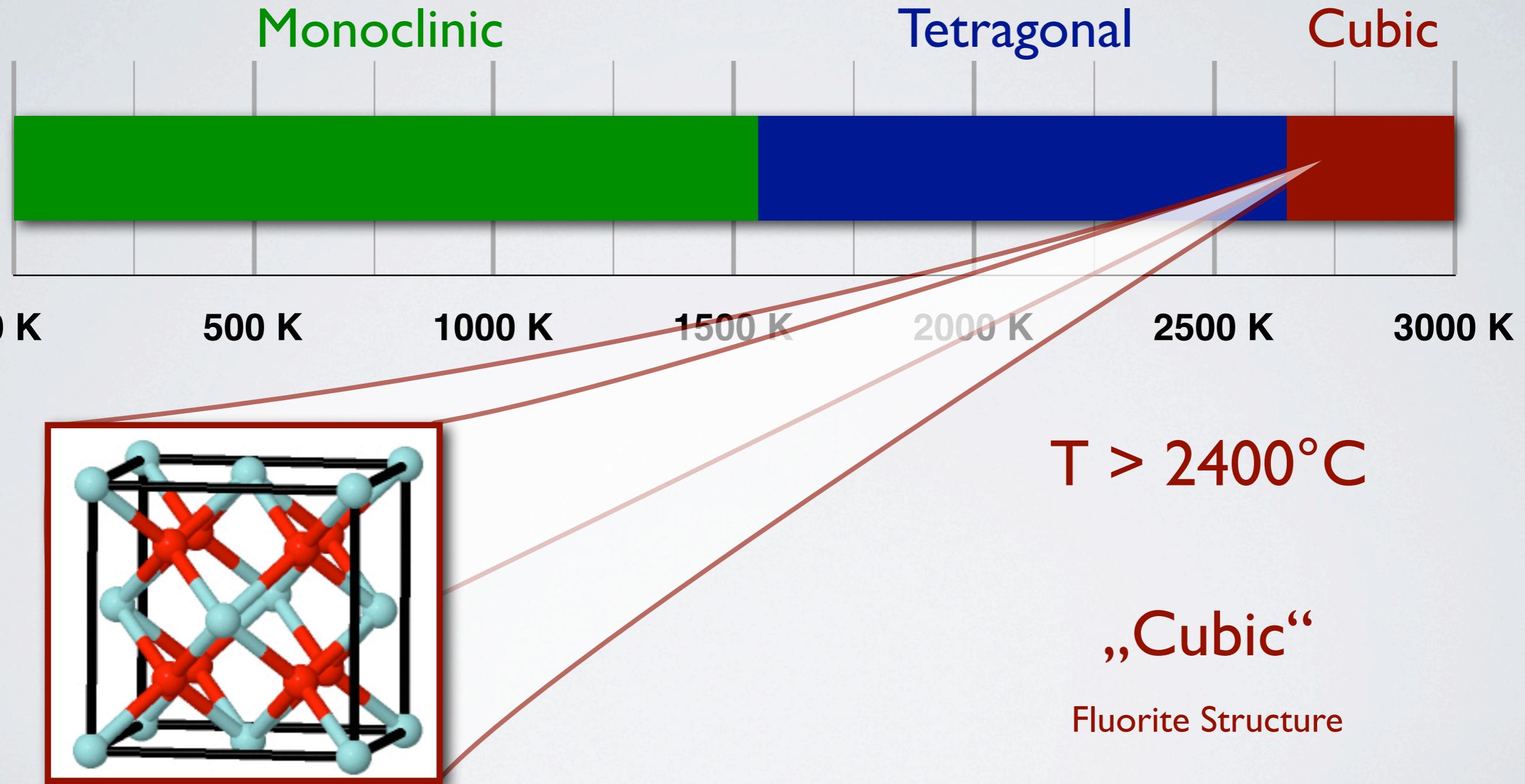
PHASE DIAGRAM OF ZrO₂



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PHASE DIAGRAM OF ZrO₂



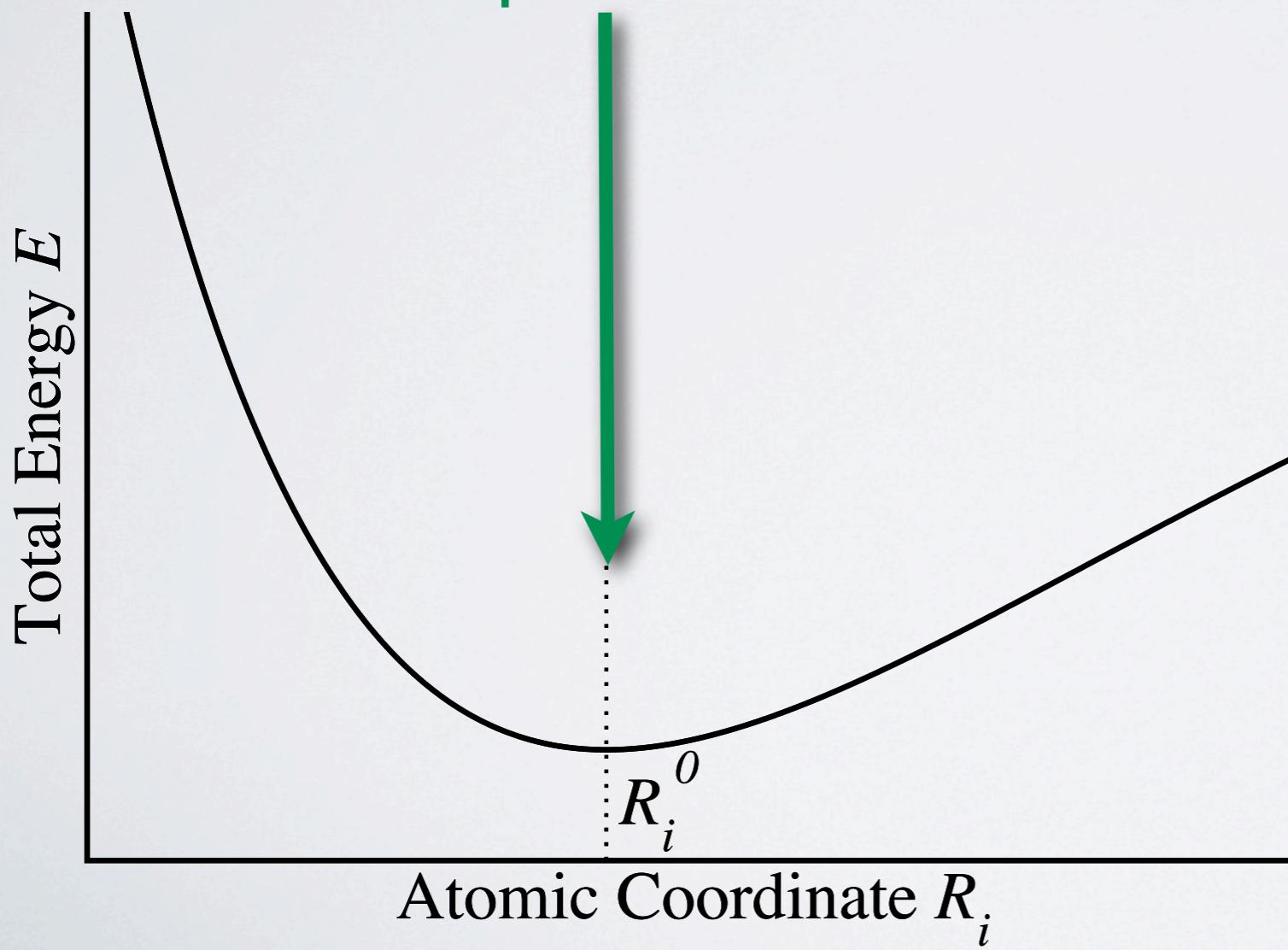
II. The Harmonic Crystal

THE INTERATOMIC INTERACTION

The total energy E is a ***3N-dimensional surface***

$$E = V(\mathbf{R}_1, \mathbf{R}_2, \dots, \mathbf{R}_N)$$

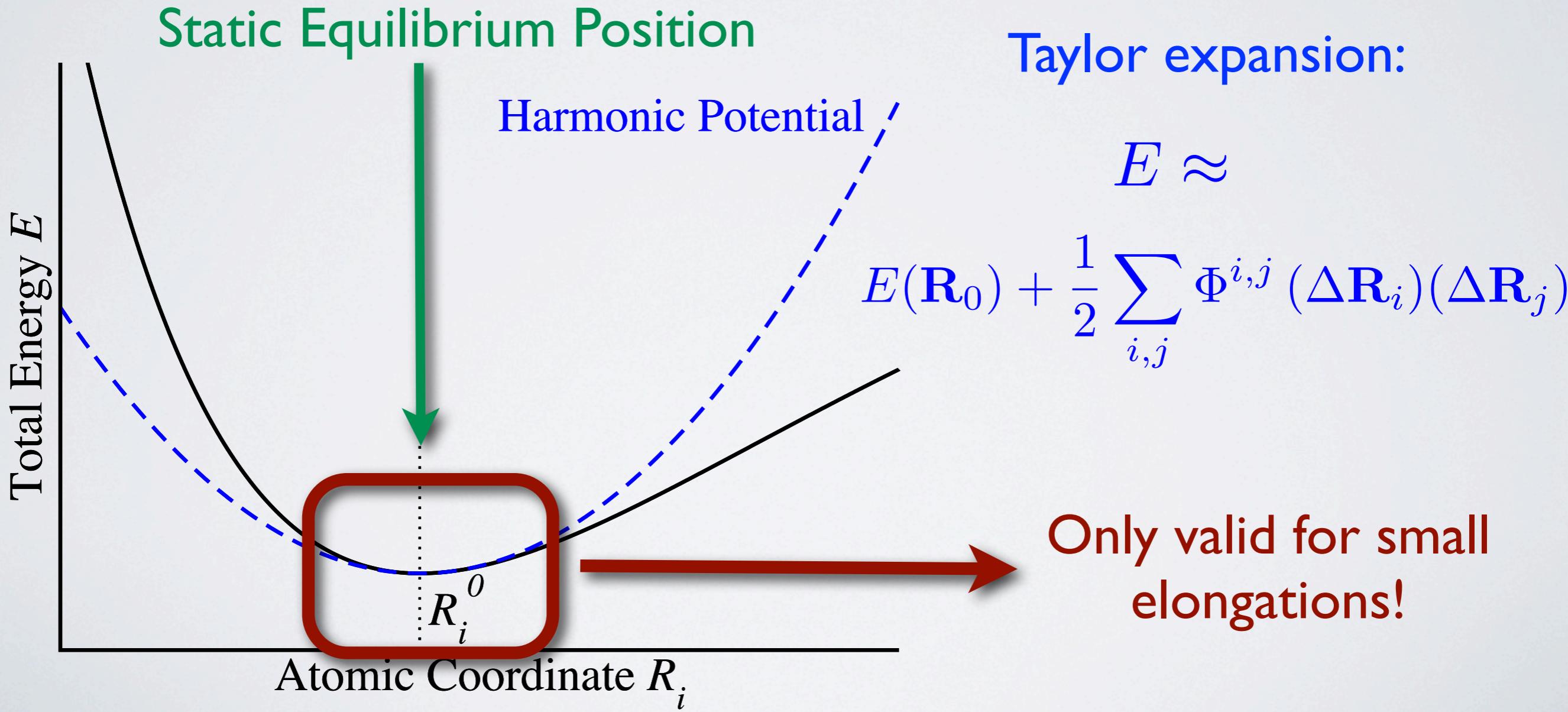
Static Equilibrium Position



THE INTERATOMIC INTERACTION

The total energy E is a **$3N$ -dimensional surface**

$$E = V(\mathbf{R}_1, \mathbf{R}_2, \dots, \mathbf{R}_N)$$



THE HARMONIC APPROXIMATION

$$E(\{\mathbf{R}_0 + \Delta\mathbf{R}\}) \approx E(\{\mathbf{R}_0\}) + \sum_i \cancel{\frac{\partial E}{\partial \mathbf{R}_i} \Big|_{\mathbf{R}_0}} \Delta\mathbf{R}_i + \frac{1}{2} \sum_{i,j} \frac{\partial^2 E}{\partial \mathbf{R}_i \partial \mathbf{R}_j} \Big|_{\mathbf{R}_0} \Delta\mathbf{R}_i \Delta\mathbf{R}_j$$

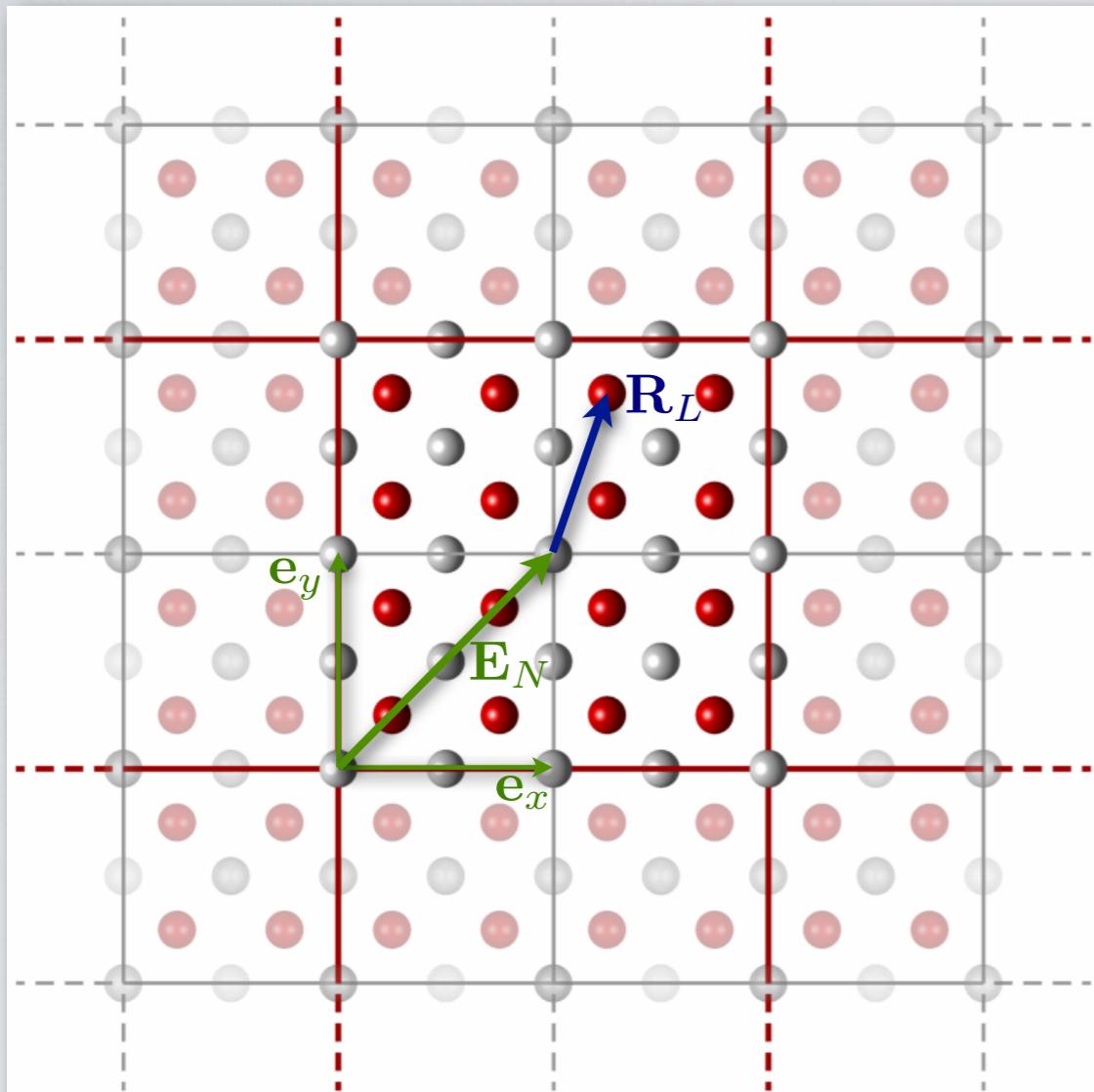
Static Equilibrium Energy Forces vanish at \mathbf{R}_0 Hessian Φ_{ij}

The diagram illustrates the harmonic approximation formula. The first term, $E(\{\mathbf{R}_0\})$, is highlighted with a green box and an arrow pointing to 'Static Equilibrium Energy'. The second term, $\sum_i \frac{\partial E}{\partial \mathbf{R}_i} \Big|_{\mathbf{R}_0} \Delta\mathbf{R}_i$, is crossed out with a large red X and an arrow pointing to 'Forces vanish at \mathbf{R}_0 '. The third term, $\frac{1}{2} \sum_{i,j} \frac{\partial^2 E}{\partial \mathbf{R}_i \partial \mathbf{R}_j} \Big|_{\mathbf{R}_0} \Delta\mathbf{R}_i \Delta\mathbf{R}_j$, is highlighted with a blue box and an arrow pointing to 'Hessian Φ_{ij} '.

Determine *harmonic force constants* Φ_{ij} :

- from **DFT-Perturbation Theory**
S. Baroni, P. Giannozzi, and A. Testa, *Phys. Rev. Lett.* **58**, 1861 (1987) &
S. Baroni, et al., *Rev. Mod. Phys.* **73**, 515 (2001).
- from **Finite Differences**
K. Kunc, and R. M. Martin, *Phys. Rev. Lett.* **48**, 406 (1982) &
K. Parlinski, Z. Q. Li, and Y. Kawazoe, *Phys. Rev. Lett.* **78**, 4063 (1997).

THE HARMONIC SOLID



Periodic Boundary Conditions
⇒ Reciprocal Space \mathbf{q}

$$D_{ij}(\mathbf{q}) = \sum_{\mathbf{E}_N} \frac{e^{i(\mathbf{q} \cdot \mathbf{E}_N)}}{\sqrt{M_i M_j}} \Phi_{ij}$$

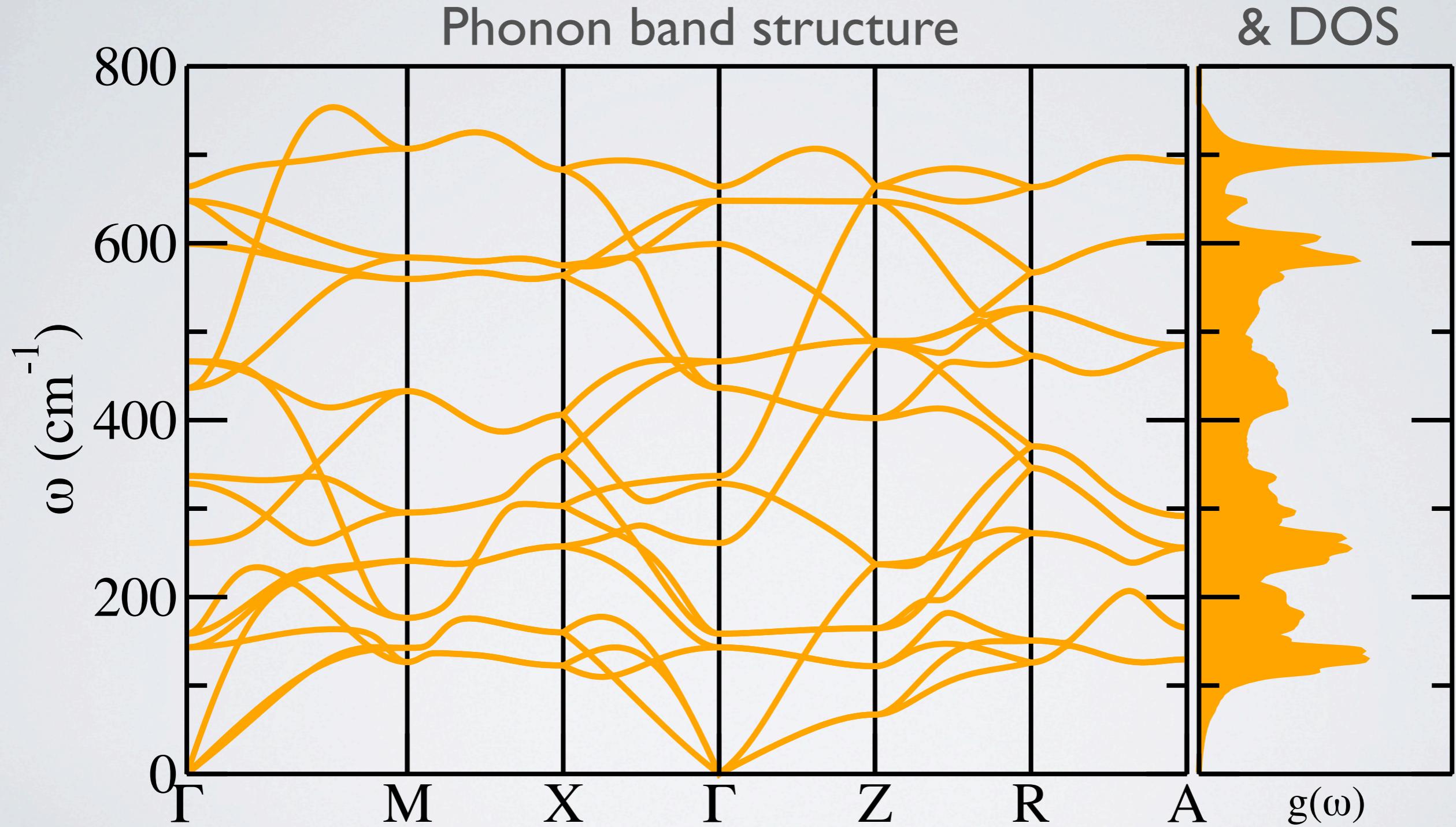
Eigenvalue problem:

$$\mathbf{D}(\mathbf{q}) [\nu(\mathbf{q})] = \omega^2(\mathbf{q}) [\nu(\mathbf{q})]$$

Real space: Superposition of harmonic oscillations

$$\mathbf{R} = R_0 + \sum_s A_s \frac{\cos(\phi_s + \omega_s t)}{\sqrt{M_i}} \cdot \nu_s$$

THE HARMONIC APPROXIMATION



THE HARMONIC FREE ENERGY

$$F^{ha}(T) = E(\{\mathbf{R}_0\})$$



Static Equilibrium Energy

$$+ \int d\omega g(\omega) \frac{\hbar\omega}{2}$$

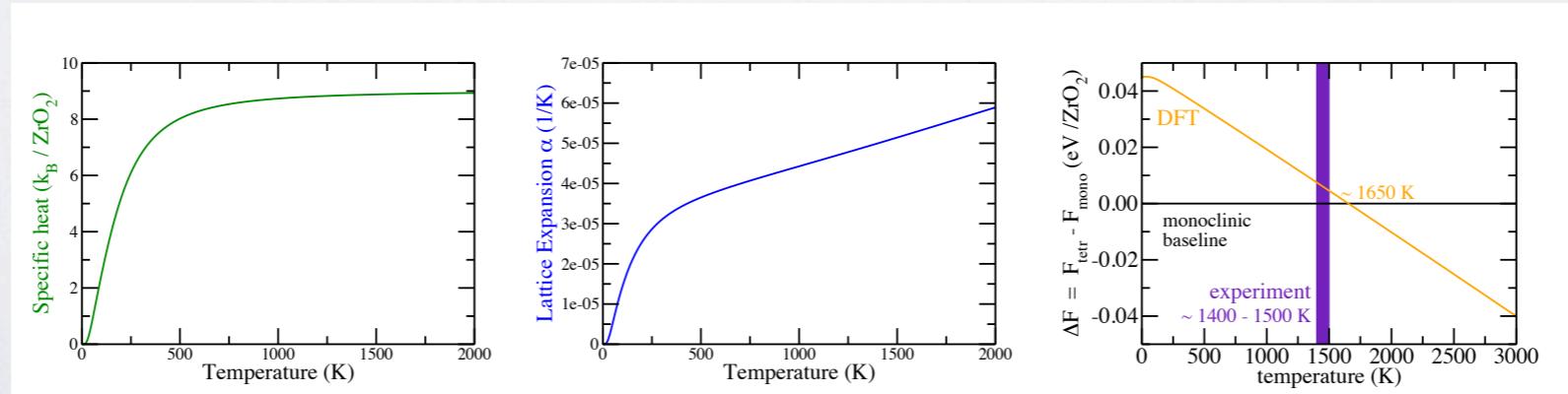


Zero-point vibration

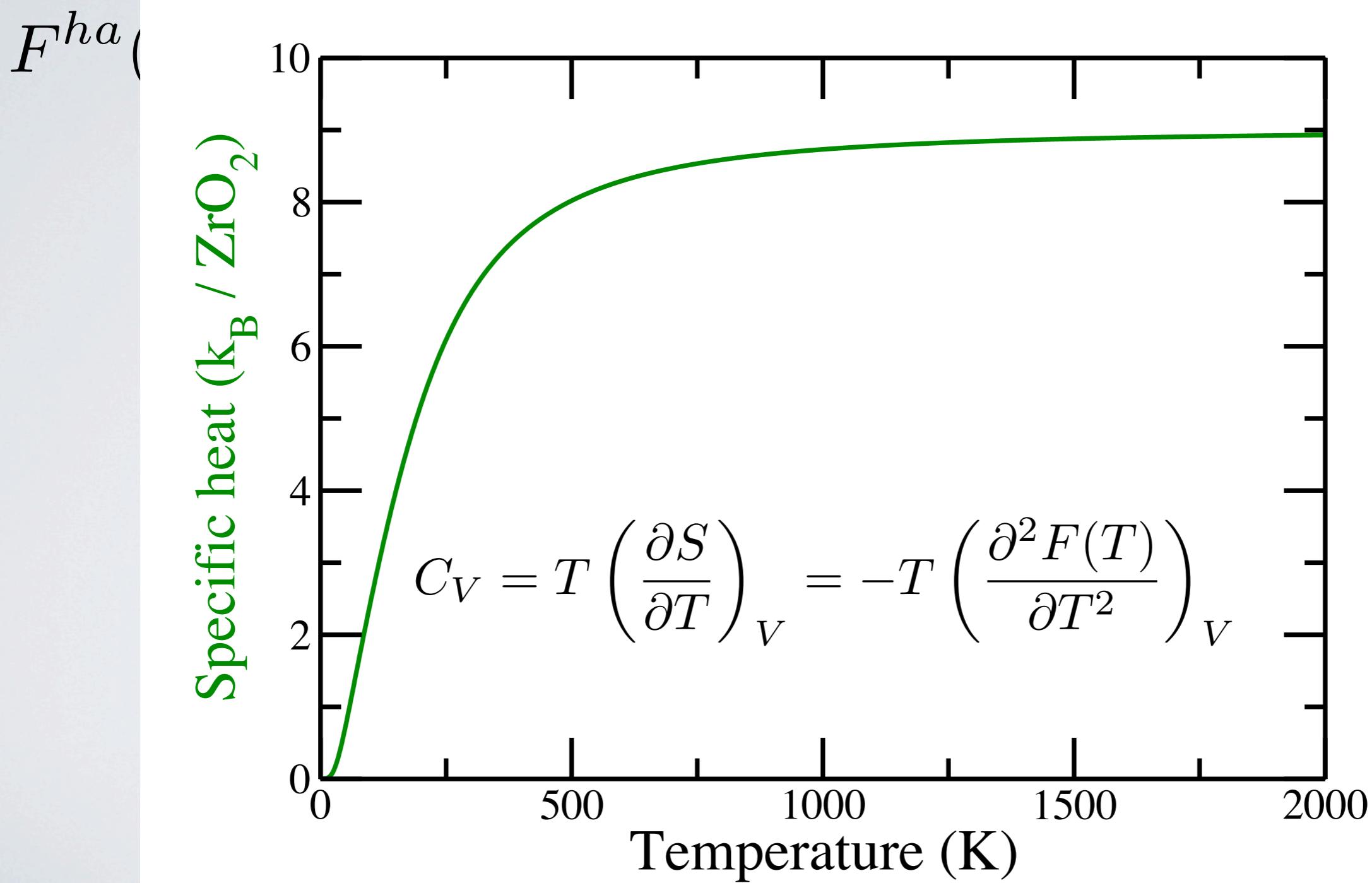
$$+ \int d\omega g(\omega) k_B T \ln \left(1 - e^{-\frac{\hbar\omega}{k_B T}} \right)$$



Thermally induced vibrations



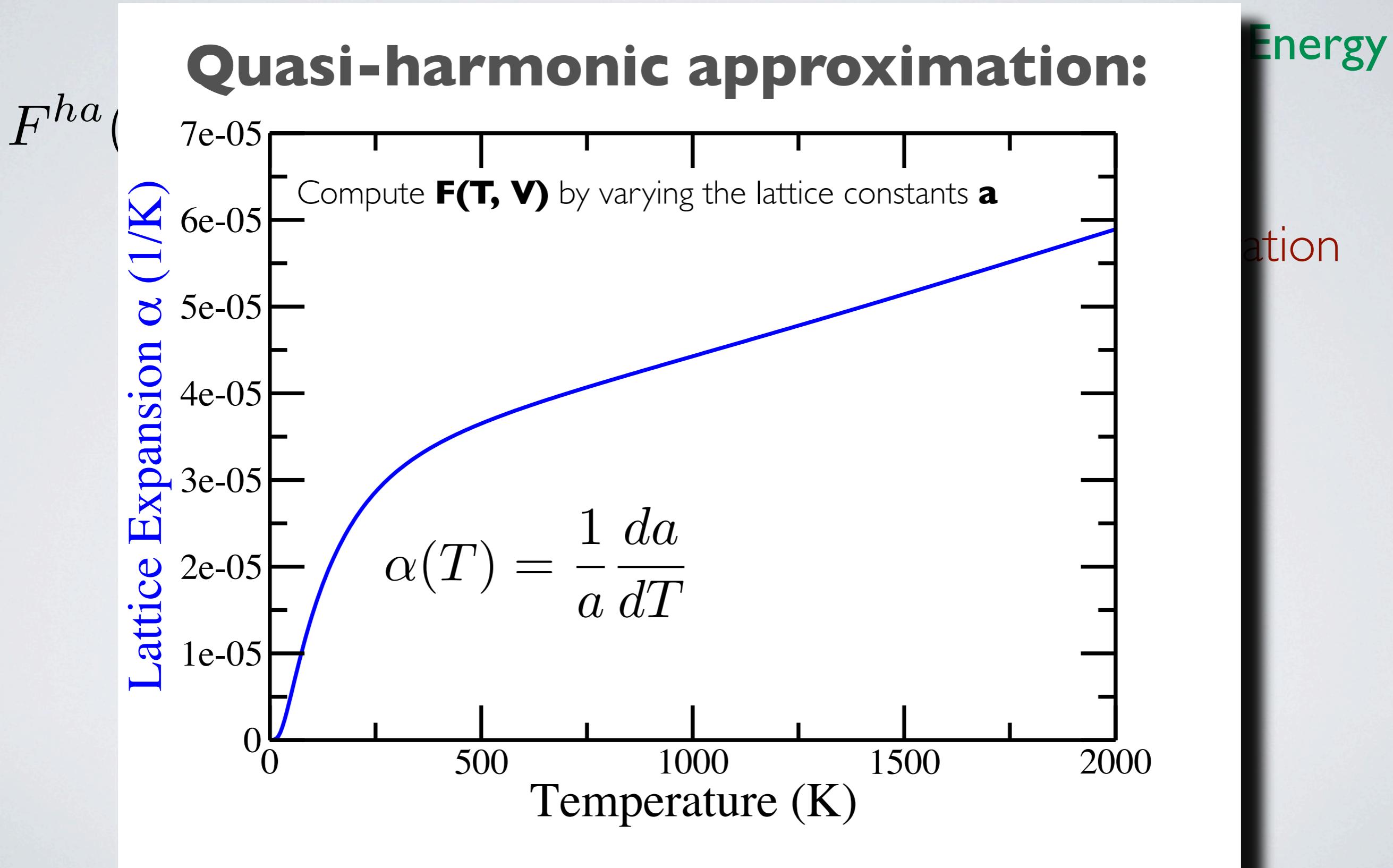
THE HARMONIC FREE ENERGY



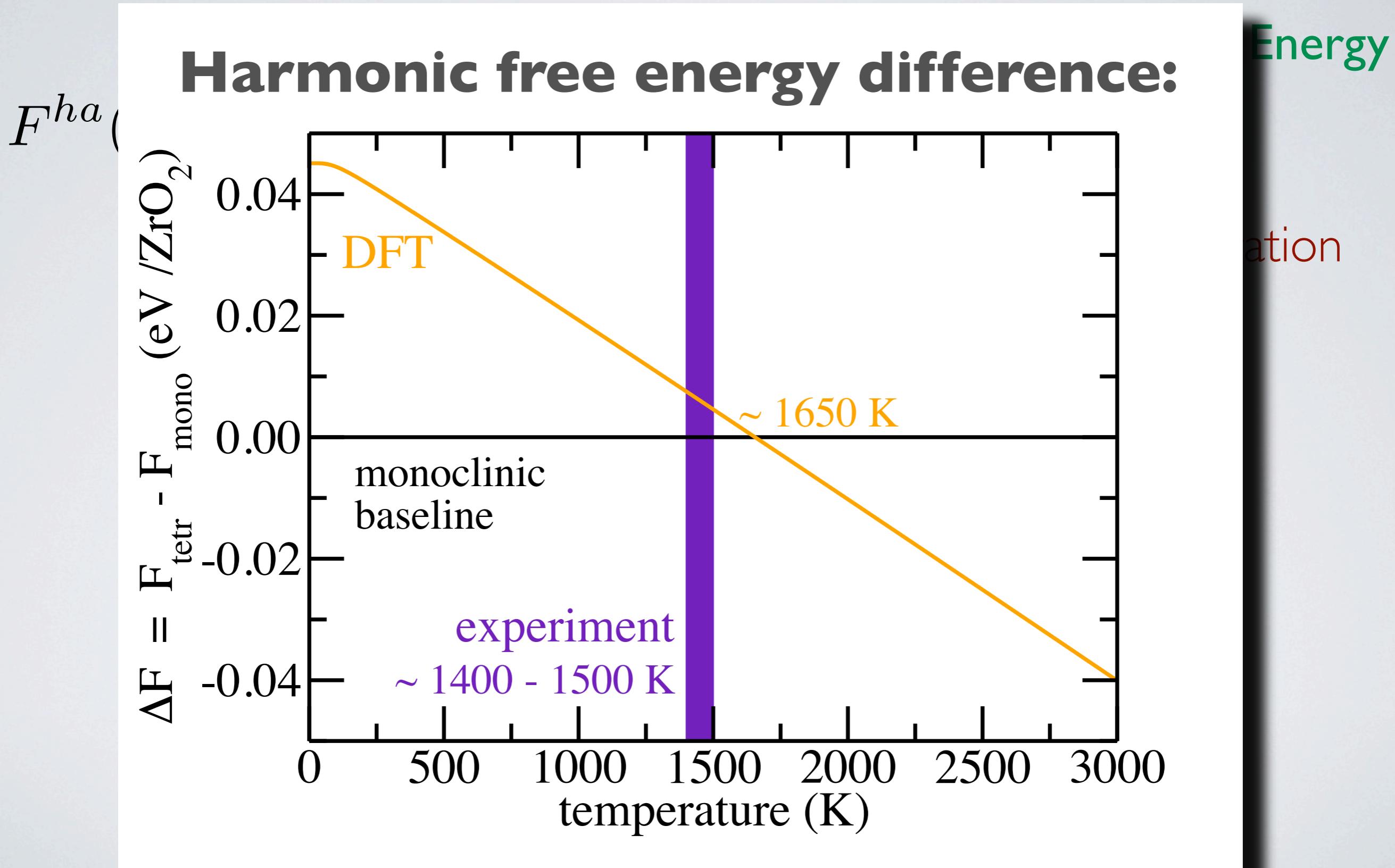
Energy

ation

THE HARMONIC FREE ENERGY

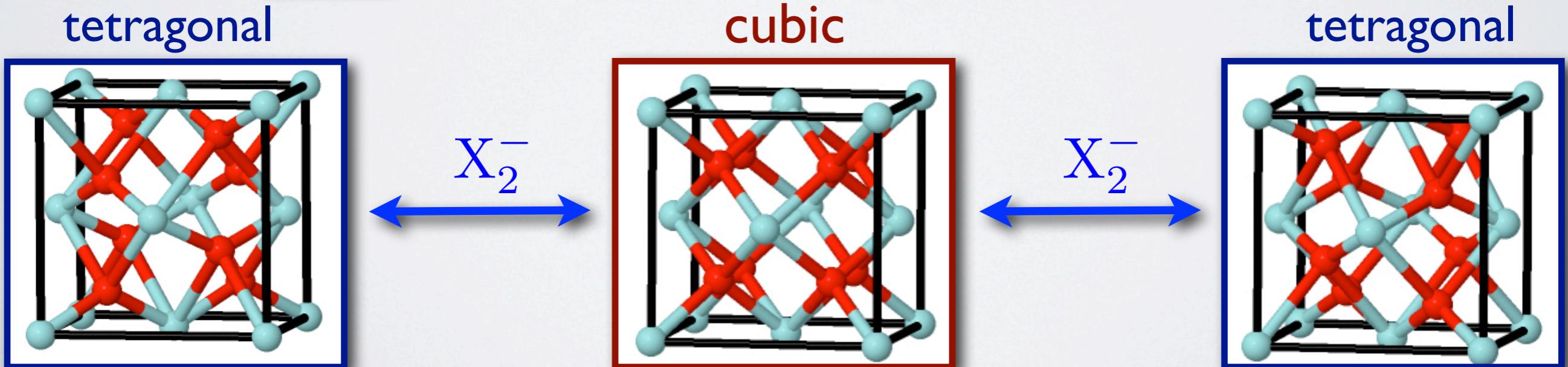
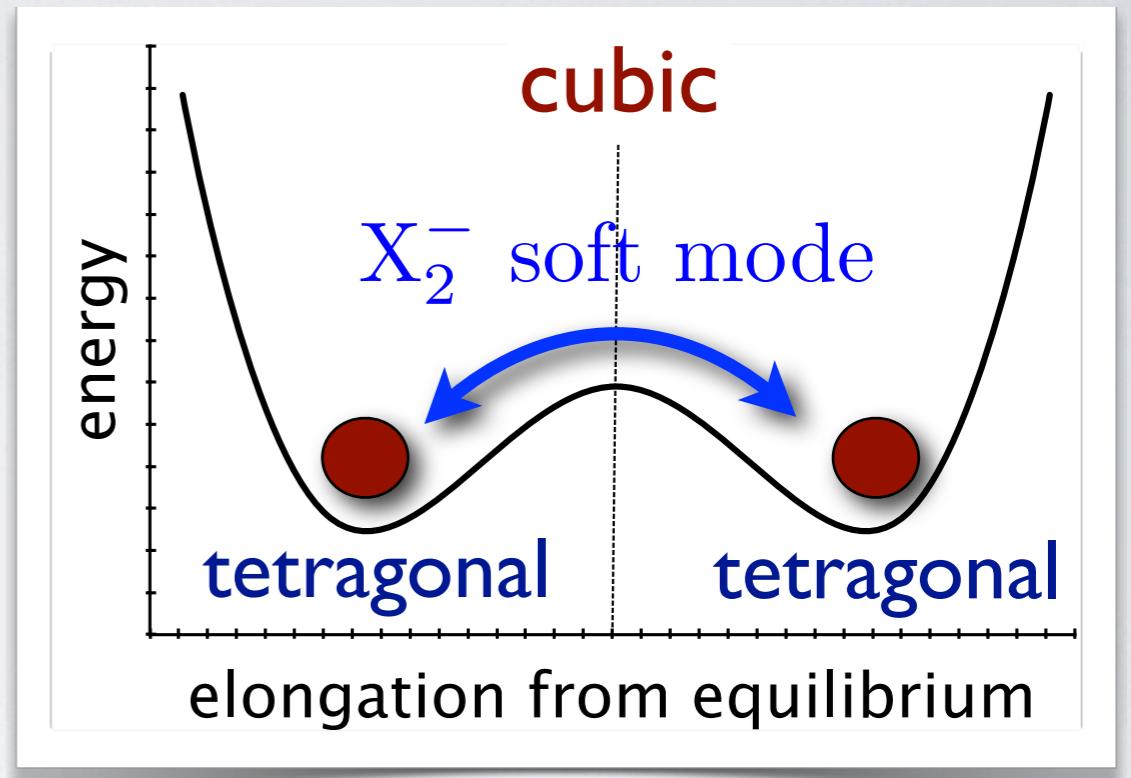
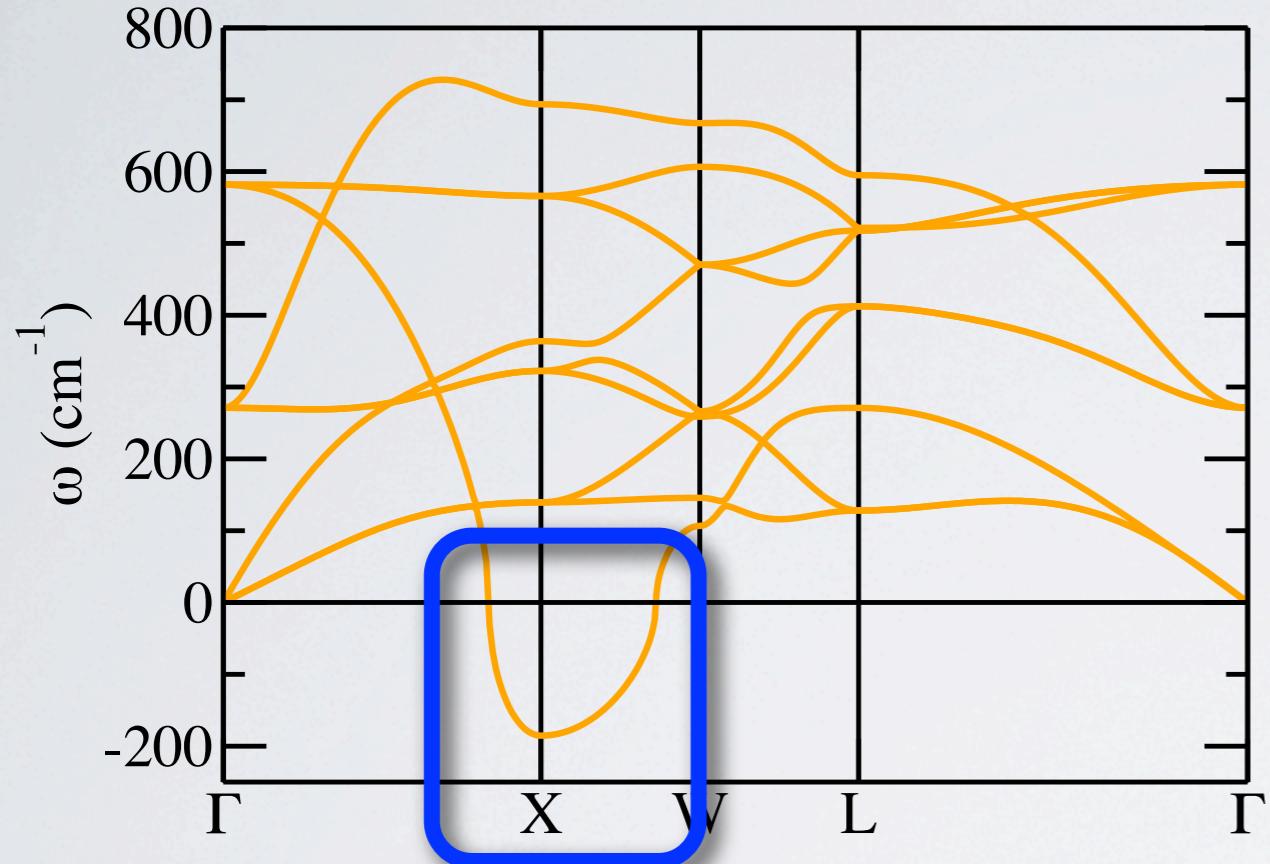


THE HARMONIC FREE ENERGY



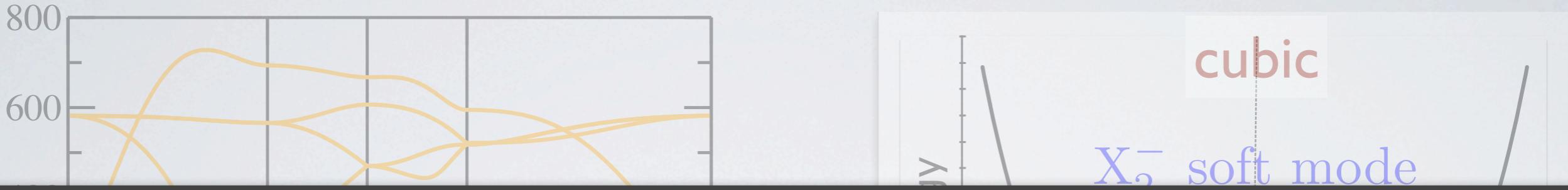
THE CUBIC ZrO_2 STRUCTURE

K. Parlinski, Z. Q. Li, and Y. Kawazoe, *Phys. Rev. Lett.* **78**, 4063 (1997).

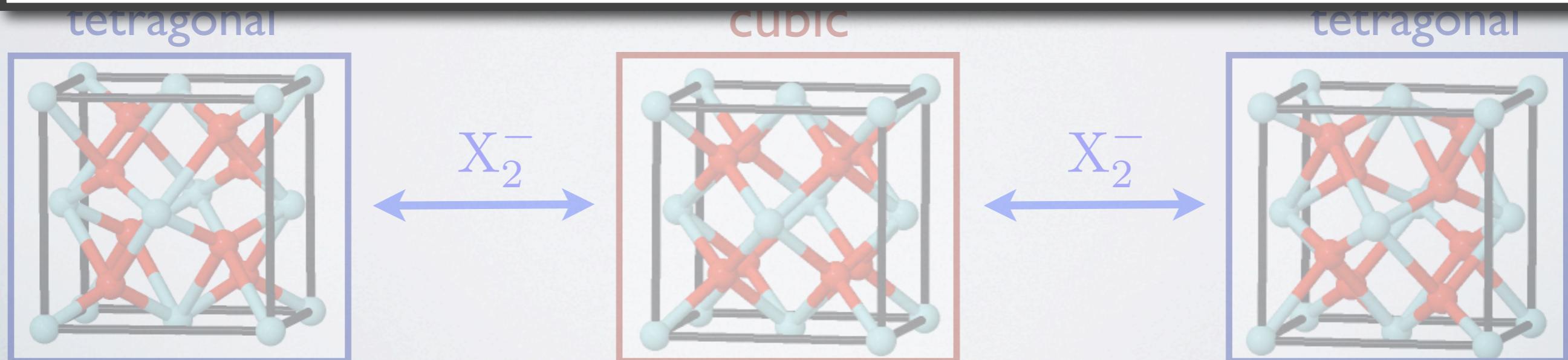


THE CUBIC ZrO_2 STRUCTURE

K. Parlinski, Z. Q. Li, and Y. Kawazoe, *Phys. Rev. Lett.* **78**, 4063 (1997).



THE HARMONIC APPROXIMATION
DOES NOT HOLD IN THE CASE
OF SOFT MODES!



THERMODYNAMIC INTEGRATION

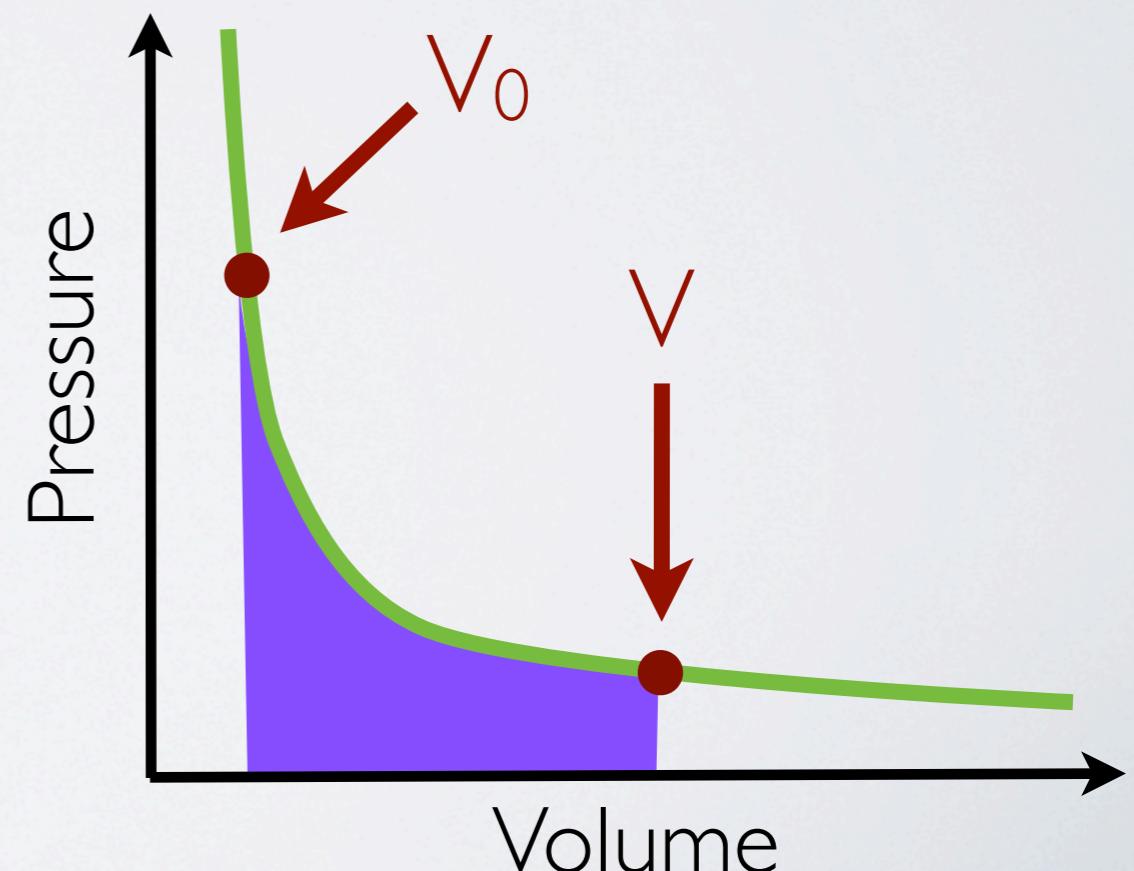
J. G. Kirkwood, *J. Chem. Phys.*, **3**, 300 (1935).

The absolute value of the **Thermodynamic Potentials** is **not measurable**, but their **derivatives** are, e.g., **the pressure**

$$p = - \left(\frac{\partial F(T, V)}{\partial V} \right)_T$$

$$F(V) - F(V_0) = - \int_{V_0}^V p \, dV$$

Free energy differences can be computed by **integration** along a **thermodynamic path**!



THERMODYNAMIC INTEGRATION

J. G. Kirkwood, *J. Chem. Phys.*, **3**, 300 (1935).

Free energy **differences** can be computed by integration along a λ -parametrized path that connects two **distinct** systems.

e.g., M. Watanabe and W. P. Reinhardt , *Phys. Rev. Lett.* **65**, 3301(1990) & O. Sugino and R. Car, *Phys. Rev. Lett.* **74**, 1823 (1995).

$$F_{\text{anh}}(T, V) = F_{\text{DFT}}(T, V) - F_{\text{har}}(T, V) = \int_0^1 d\lambda \left\langle \frac{\partial U_{\text{hyb}}(\lambda)}{\partial \lambda} \right\rangle_{\text{hyb}}$$

THERMODYNAMIC INTEGRATION

J. G. Kirkwood, *J. Chem. Phys.*, **3**, 300 (1935).

Free energy **differences** can be computed by integration along a λ -parametrized path that connects two **distinct** systems.

e.g., M. Watanabe and W. P. Reinhardt , *Phys. Rev. Lett.* **65**, 3301(1990) & O. Sugino and R. Car, *Phys. Rev. Lett.* **74**, 1823 (1995).

„Complete“ free energy Harmonic reference
free energy

$$F_{\text{anh}}(T, V) = F_{\text{DFT}}(T, V) - F_{\text{har}}(T, V) = \int_0^1 d\lambda \left\langle \frac{\partial U_{\text{hyb}}(\lambda)}{\partial \lambda} \right\rangle_{\text{hyb}}$$

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Thermodynamic expectation value for the „hybrid“ system

$$U_{\text{hyb}}(\lambda) = \lambda \cdot U_{\text{DFT}} + (1 - \lambda) \cdot U_{\text{har}}$$

ab initio potential *harmonic potential*

THERMODYNAMIC INTEGRATION

J. G. Kirkwood, *J. Chem. Phys.*, **3**, 300 (1935).

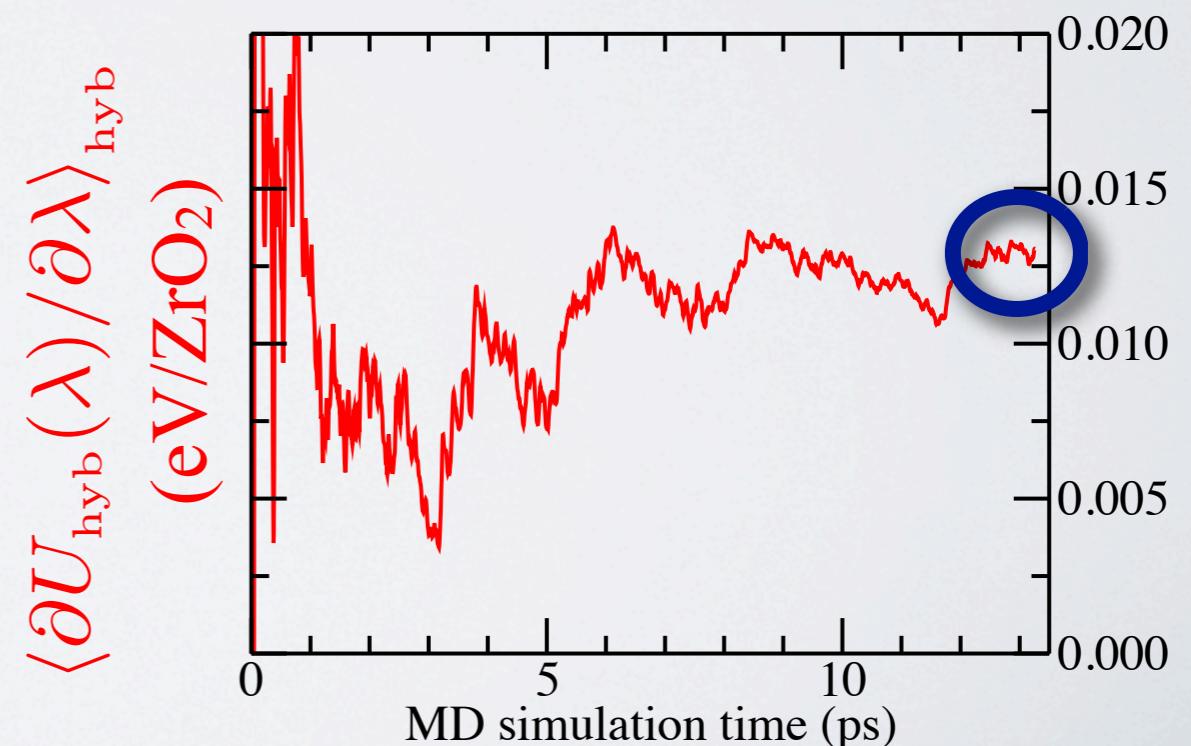
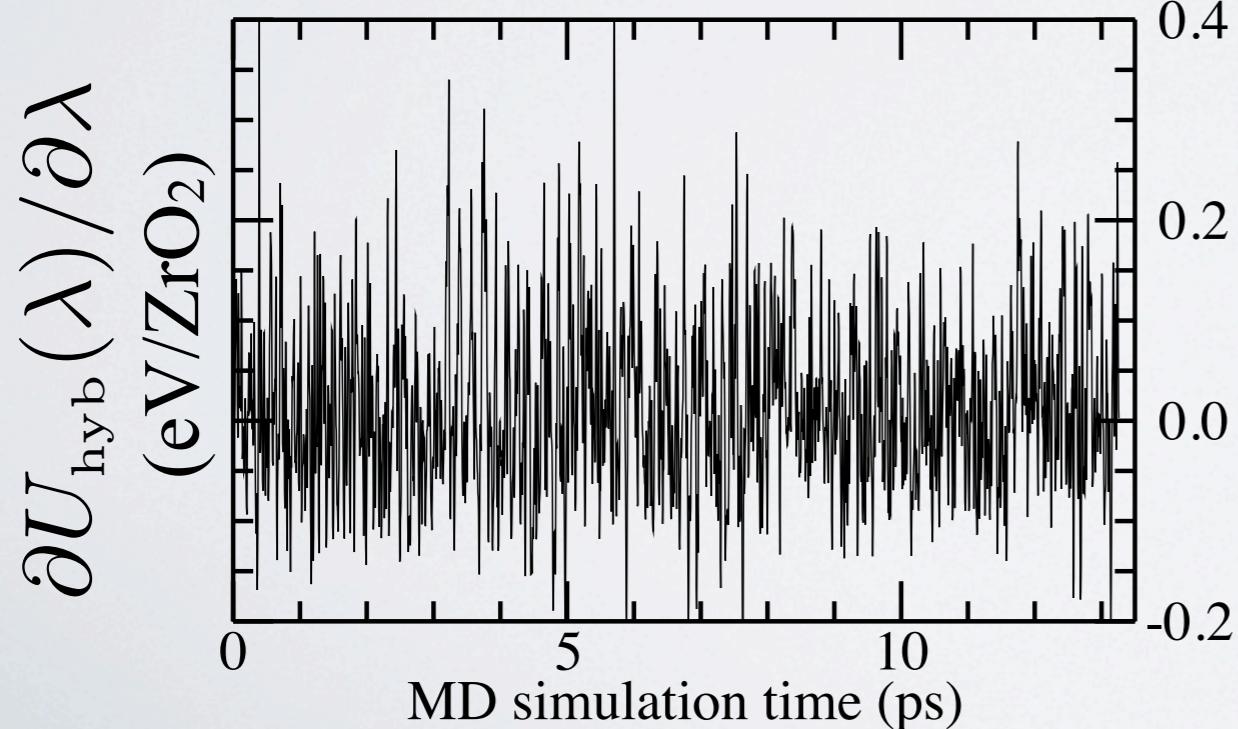
Molecular Dynamics for the hybrid system:

B. Grabowski et al., *Phys. Rev. B* **79**, 134106 (2009).

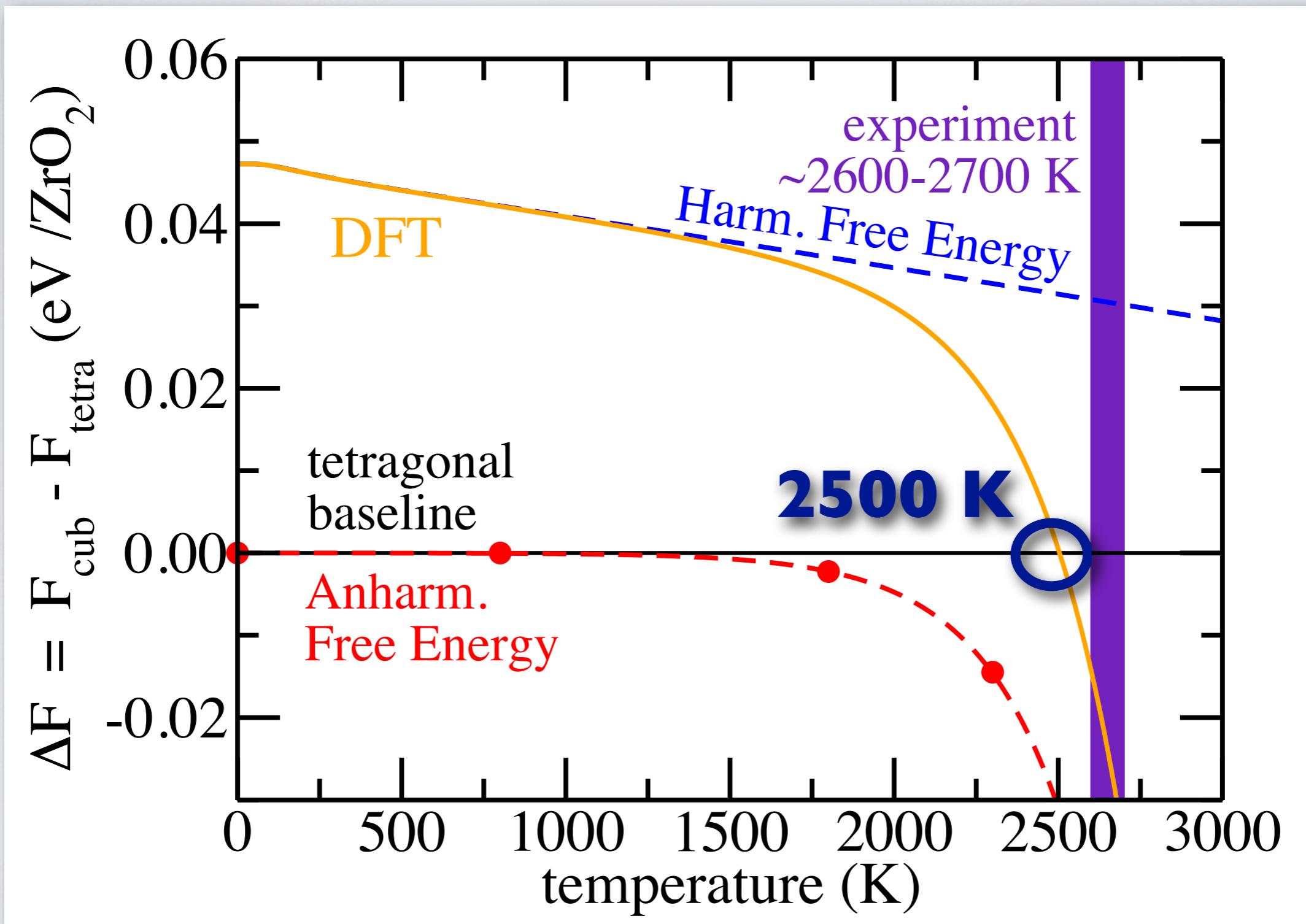
$$U_{\text{hyb}}(\lambda) = \lambda \cdot U_{\text{DFT}} + (1 - \lambda) \cdot U_{\text{har}}$$

ab initio
potential harmonic potential

LDA, $\lambda = 0.5$, $T = 2800$ K



CUBIC PHASE STABILITY

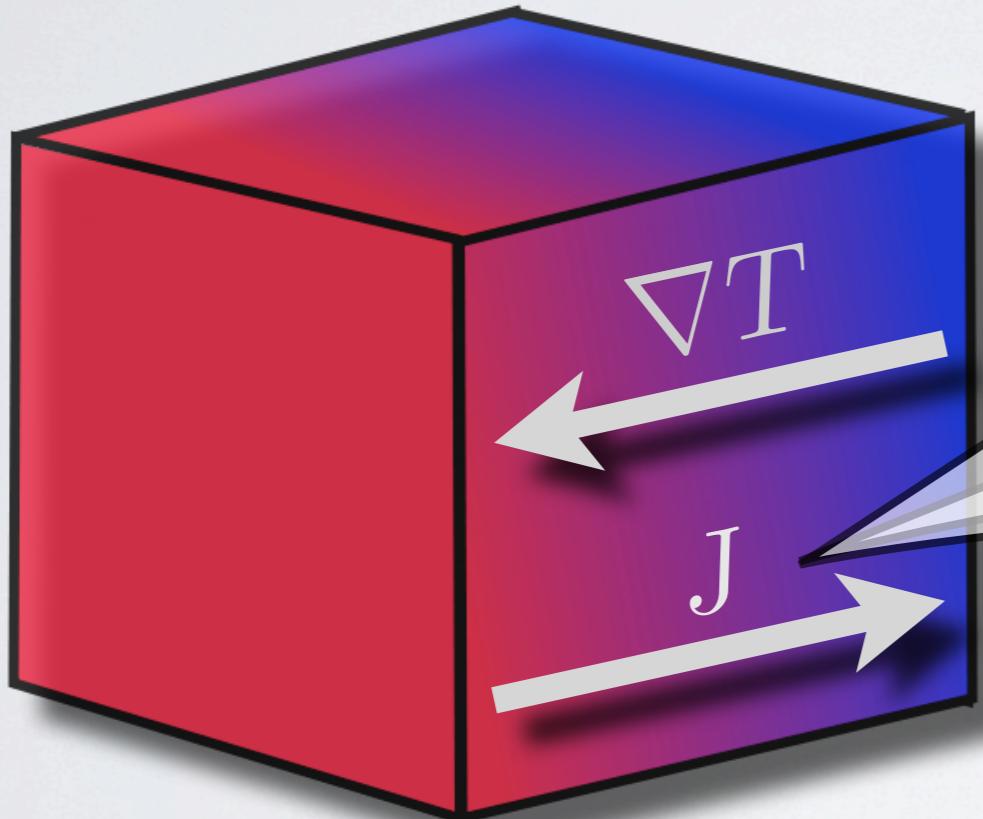


Noticeable Anharmonic Effects for $T > 1500$ K !

III. Non-Equilibrium Thermodynamics

THERMAL CONDUCTIVITY

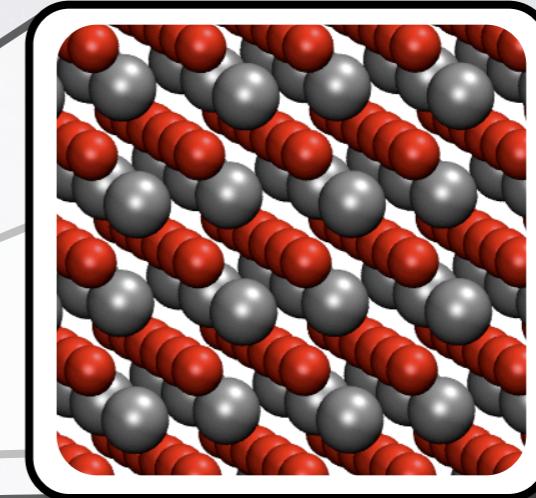
**Macroscopic
Effect:**



Fourier's Law:

$$\mathbf{J} = -\kappa \nabla T$$

**Microscopic
Nature:**

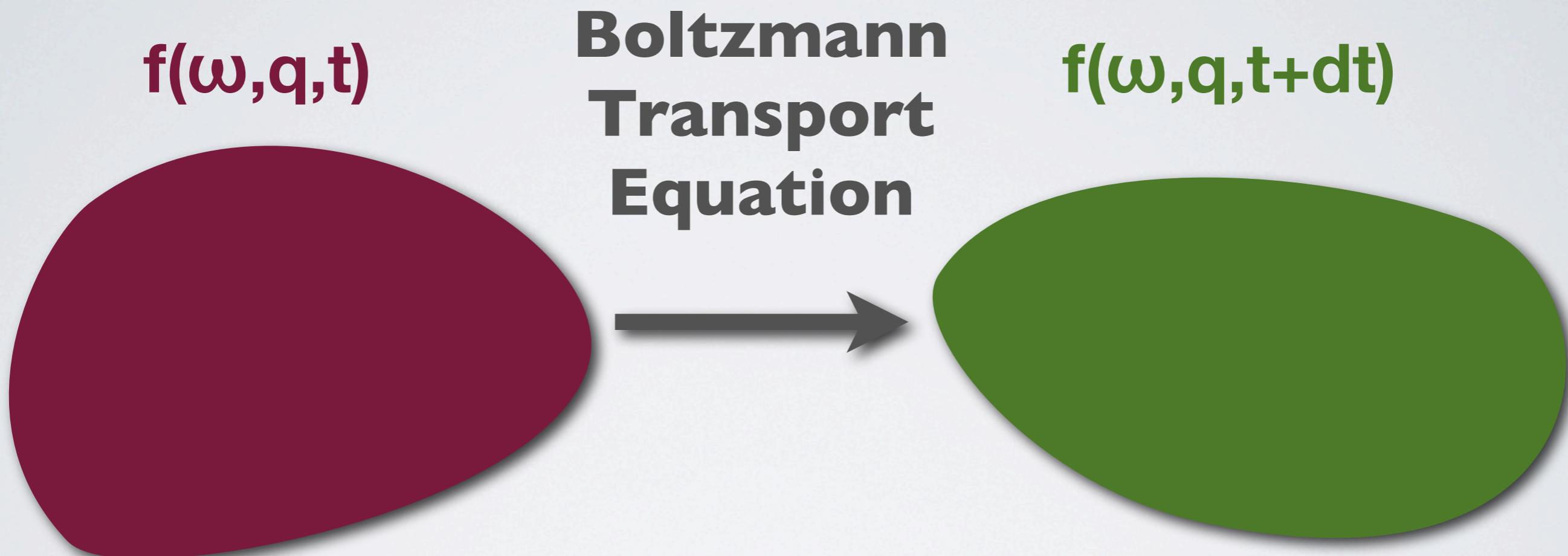


Can we calculate and understand the **thermal conductivity** from **first principles**?

(A) BOLTZMANN TRANSPORT EQUATION

R. Peierls, *Ann. Phys.* **395**, 1055 (1929).

D.A. Broido et al., *Appl. Phys. Lett.* **91**, 231922 (2007).



Boltzmann-Peierls-Transport-Equation describes the evolution of the **phonon** phase space distribution $f(\omega, q, t)$.

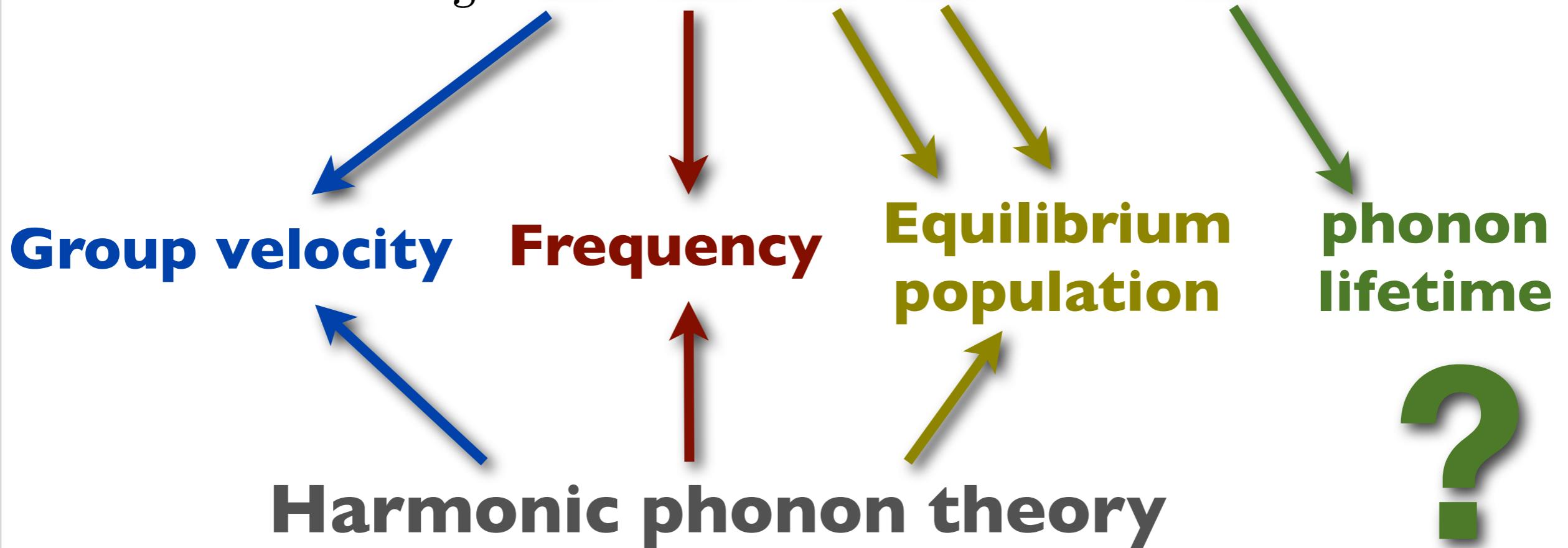
(A) BOLTZMANN TRANSPORT EQUATION

R. Peierls, Ann. Phys. **395**, 1055 (1929).

D.A. Broido et al., Appl. Phys. Lett. **91**, 231922 (2007).

Single-mode relaxation time approximation

$$\kappa \sim \sum_s c_s^2 \omega_s^2 n_s (n_s + 1) \tau_s$$



Phonon Lifetimes from First Principles

- from **Density Functional Perturbation Theory**

J. Garg et al., *Phys. Rev. Lett.* **106**, 045901 (2011).

- from **fitting the forces** in *ab initio MD*

K. Esfarjani, and H.T. Stokes, *Phys. Rev. B* **77**, 144112 (2008).

- from **fitting the phonon line width** determined via *ab initio MD*

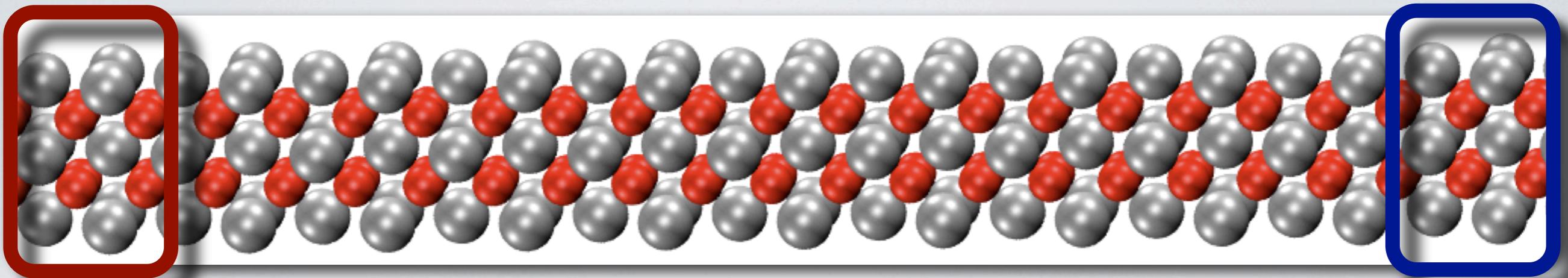
N. De Koker, *Phys. Rev. Lett.* **103**, 125902 (2009).

All these approaches give very **accurate** results for **good thermal conductors** at **low** temperatures.

Results are **questionable** at high levels of **anharmonicity**!

(B) NON-EQUILIBRIUM MD

S. Stackhouse, L. Stixrude, and B. B. Karki, *Phys. Rev. Lett.* **104**, 208501 (2010).



**heat
source**



- Temperature gradient ∇T
- Stationary heat flux J



**heat
sink**



Thermal conductivity can be calculated
by applying Fourier's Law.

$$J = -\kappa \nabla T$$

PROS & CONS

- Relatively easy implementation:**
Use standard thermostats for heat source & sink
- Modeling of the steady state**
Long simulation time needed for good statistics
- Heat source & sink create artifacts**
Large simulation cells needed
- Huge temperature gradient**
Possible undesired non-linear effects

FLUCTUATION-DISSIPATION THEOREM

Brownian Motion:

A. Einstein, Ann. Phys. **322**, 549 (1905).

The erratic motion of the particles
is closely related to
frictional force under perturbation.



The fluctuations of the forces in thermodynamic equilibrium is related to the generalized resistance in non-equilibrium for linear dissipative systems.

H. B. Callen, and T.A. Welton, Phys. Rev. **83**, 34 (1951).

(C) GREEN-KUBO METHOD

R. Kubo, M. Yokota, and S. Nakajima, *J. Phys. Soc. Japan* **12**, 1203 (1957).

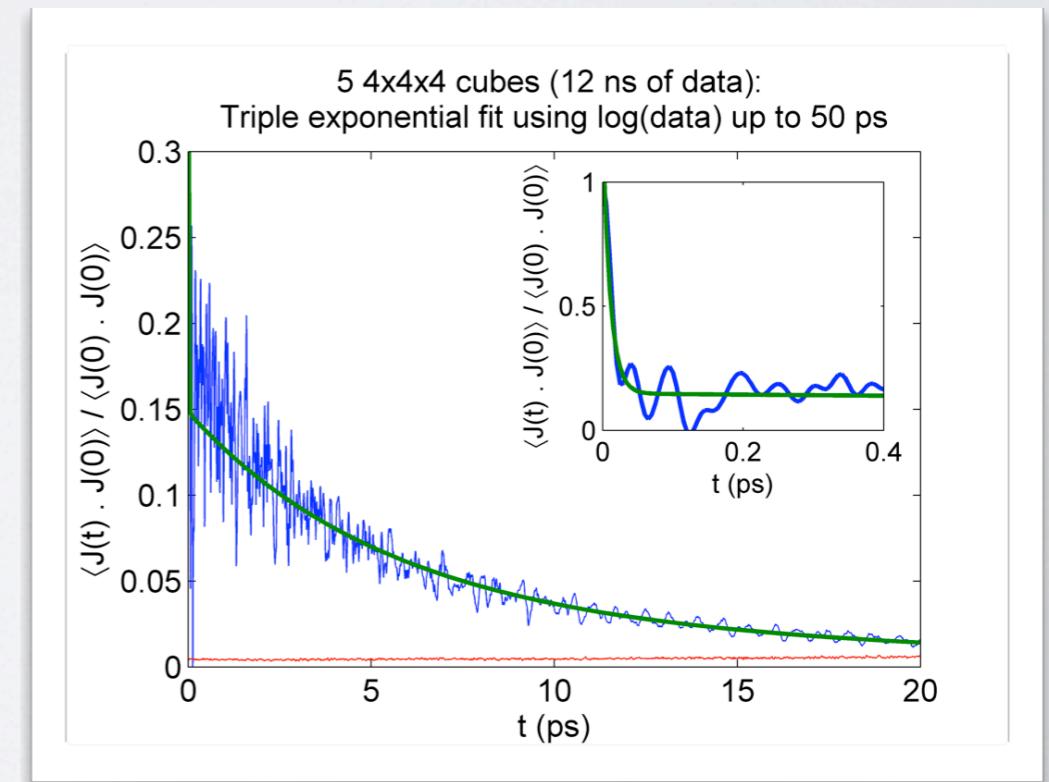
Simulations of the **thermodynamic equilibrium**



Information about **non-equilibrium processes**

$$\kappa \sim \int_0^\infty d\tau \langle \mathbf{J}(0) \cdot \mathbf{J}(\tau) \rangle_{eq}$$

The **thermal conductivity** is related to the **autocorrelation function** of the **heat flux**



PROS & CONS

- Standard MD method
- No heat source & sink needed
- No temperature gradient needed
- Long simulation time

Looks promising, but...

THE HEAT FLUX

R. J. Hardy, Phys. Rev. **132**, 168 (1963).

$$\mathbf{J}(t) = \frac{d}{dt} \left(\sum_i \mathbf{r}_i(t) \boxed{\varepsilon_i(t)} \right)$$

$\mathbf{r}_i \dots$ Position of atom i
 $\varepsilon_i \dots$ Energy of atom i

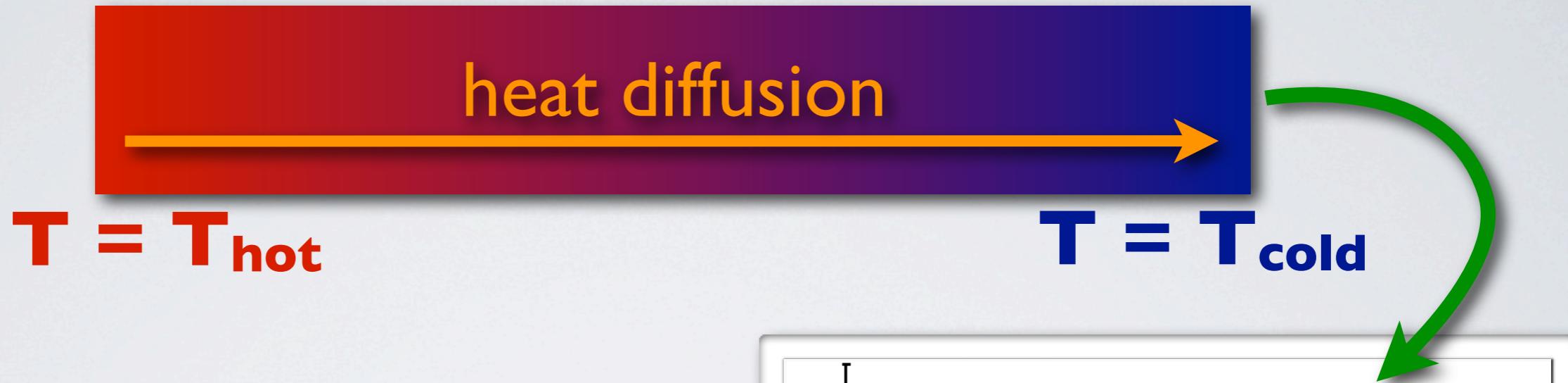
Energetic contributions ε_i of the **single atoms** required!

A partitioning of the **total energy** is ill-defined, cumbersome and error-prone!

⇒ **Green-Kubo Method hitherto only used with classical potentials!**

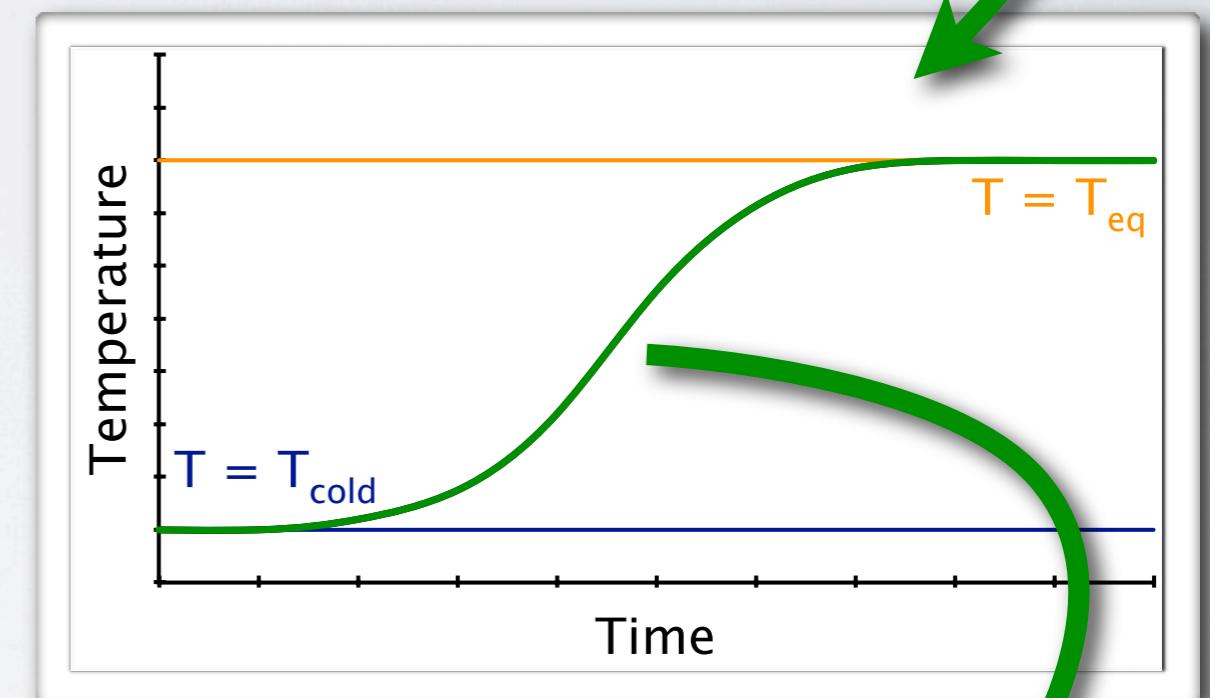
„LASER FLASH“ MEASUREMENTS

W. J. Parker et al., J. Appl. Phys. **32**, 1679 (1961).



Heat Diffusion Equation:

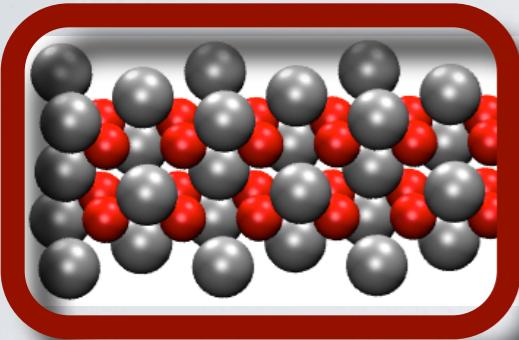
$$\frac{\partial T(x, t)}{\partial t} + \frac{\kappa}{\rho \cdot c_V} \frac{\partial^2 T(x, t)}{\partial x^2} = 0$$



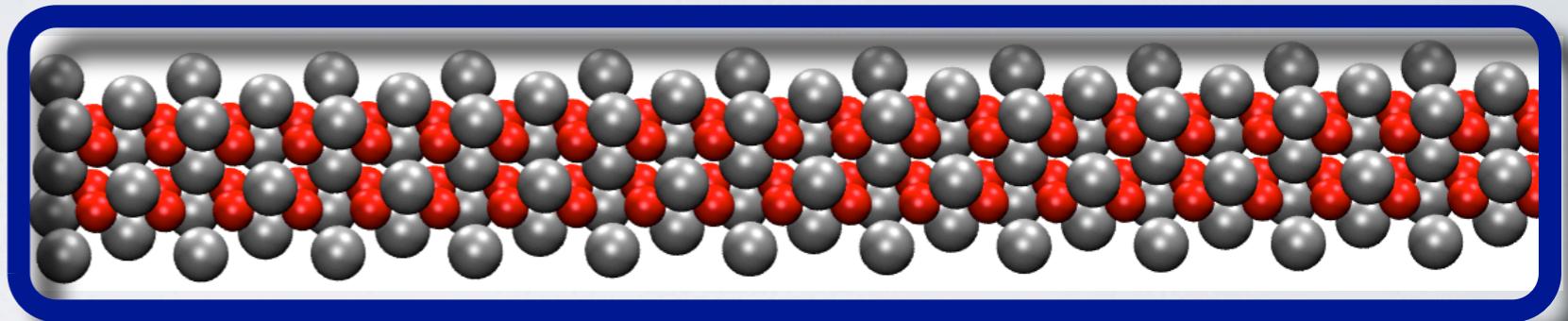
Extract the **thermal conductivity** by fitting $T(x, t)$

(D) LASER FLASH METHOD

T. M. Gibbons and S. K. Estreicher, *Phys. Rev. Lett.* **102**, 255502 (2009).



hot SC

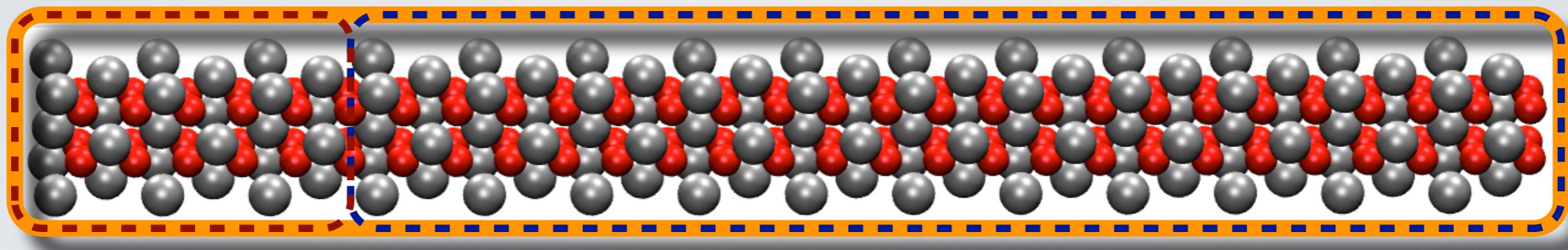


cold supercell

- (I) Prepare two supercells:
a **small hot** one and a **large cold** one.

(D) LASER FLASH METHOD

T. M. Gibbons and S. K. Estreicher, *Phys. Rev. Lett.* **102**, 255502 (2009).



Thermal Diffusion

- (1) Prepare two supercells:
a **small hot** one and a **large cold** one.
- (2) Connect the **two** supercells and
let the heat diffuse.

(I) SUPERCELL PREPARATION

S. K. Estreicher, and T. M. Gibbons, *Physica B* **404**, 4509 (2009).

In the **quasi-harmonic approximation**, the positions \mathbf{r}_i and the velocities \mathbf{v}_i are related to the vibrational eigenfrequencies ω_s and -vectors \mathbf{v}_s .

$$R_{0i} + \Delta\mathbf{R}_i = + \sum_s A_s(T) \frac{\cos(\phi_s + \omega_s t)}{\sqrt{M_i}} \cdot \nu_s$$
$$\mathbf{v}_i = - \sum_s A_s(T) \frac{\sin(\phi_s + \omega_s t)}{\sqrt{M_i}} \cdot \omega_s \cdot \nu_s$$

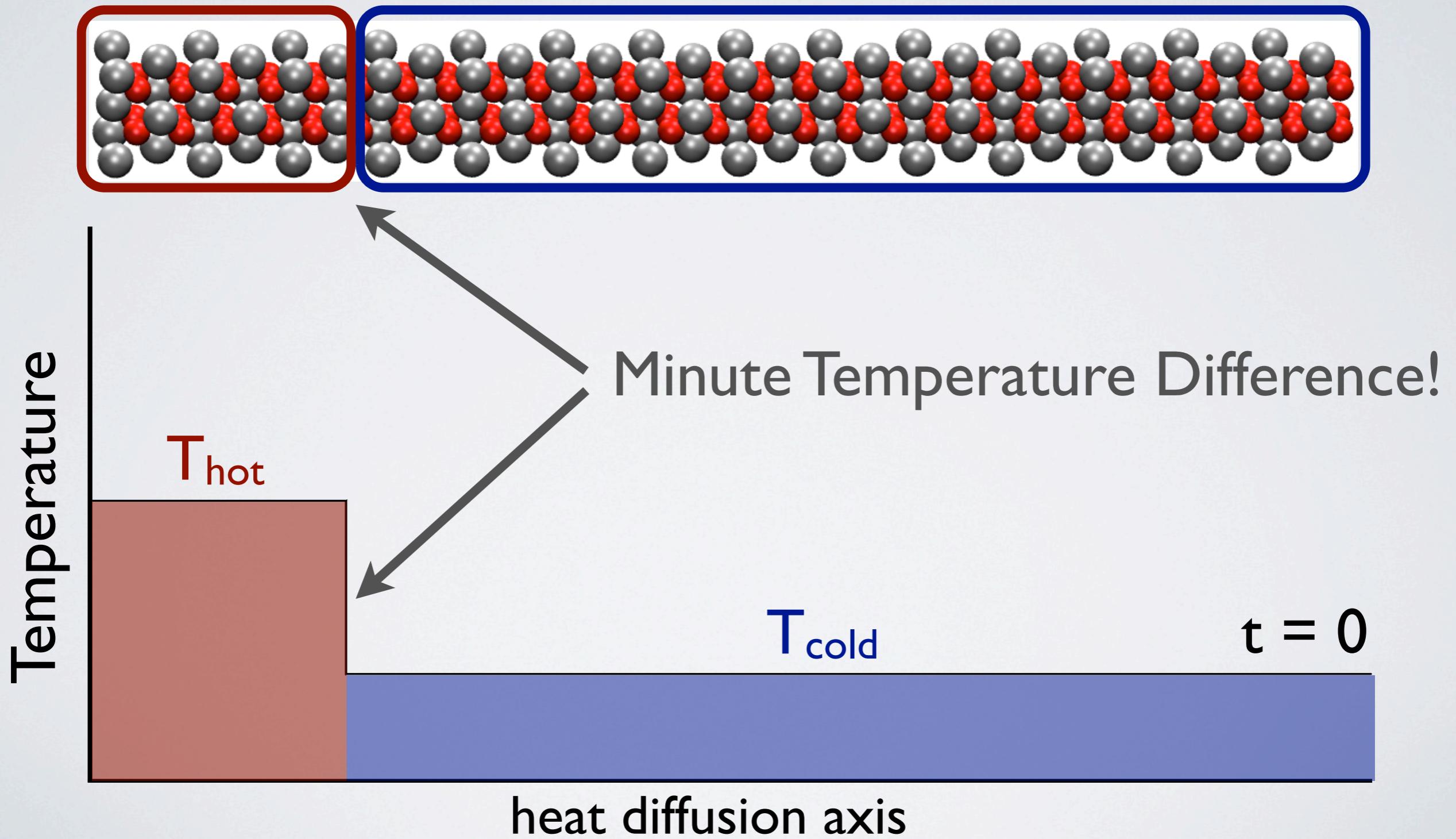
Maxwell-Boltzmann distributed amplitudes

random phase

harmonic approximation

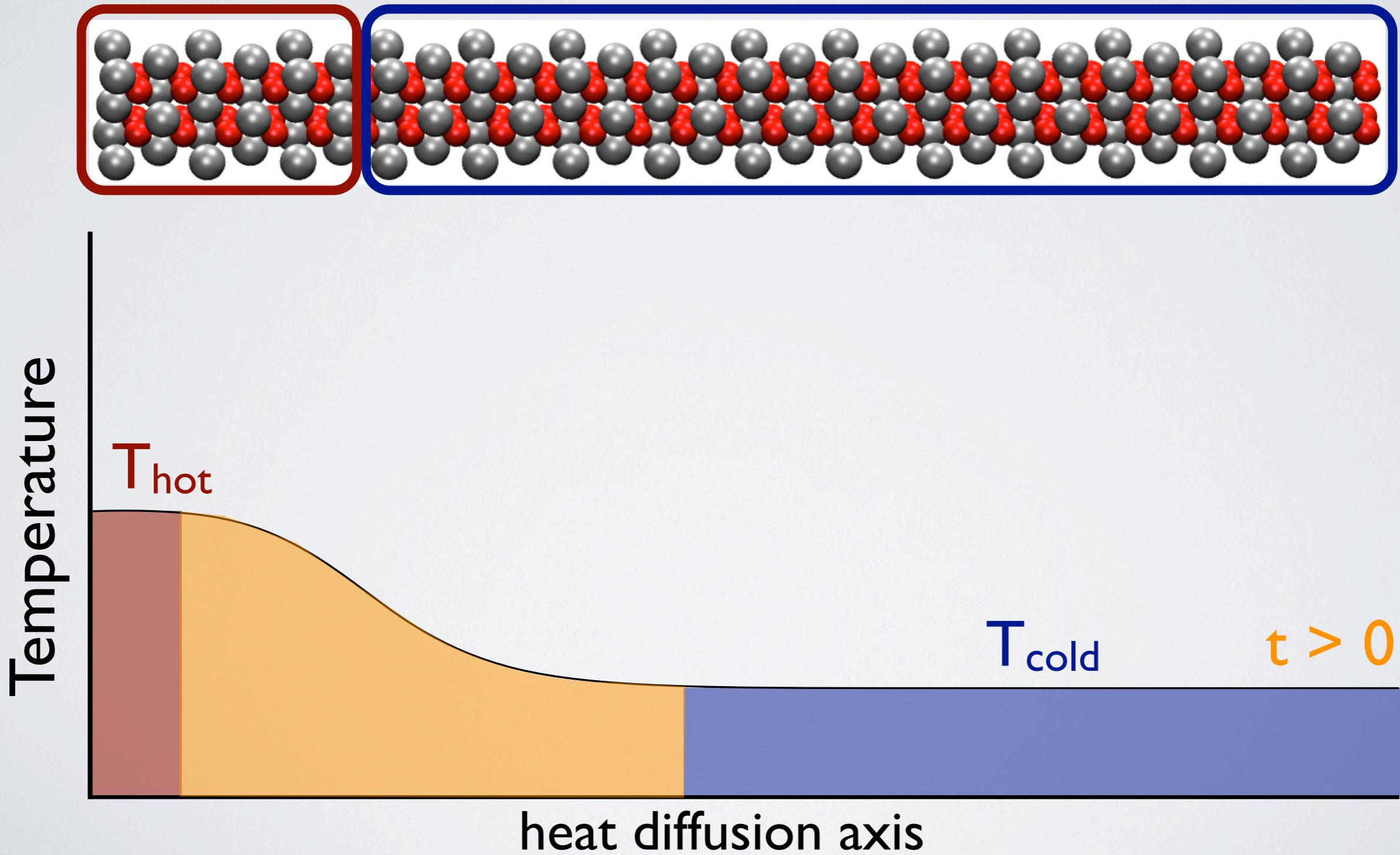
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S. K. Estreicher, and T. M. Gibbons, *Physica B* **404**, 4509 (2009).



(2) HEAT DIFFUSION

T. M. Gibbons and S. K. Estreicher, Phys. Rev. Lett. **102**, 255502 (2009).

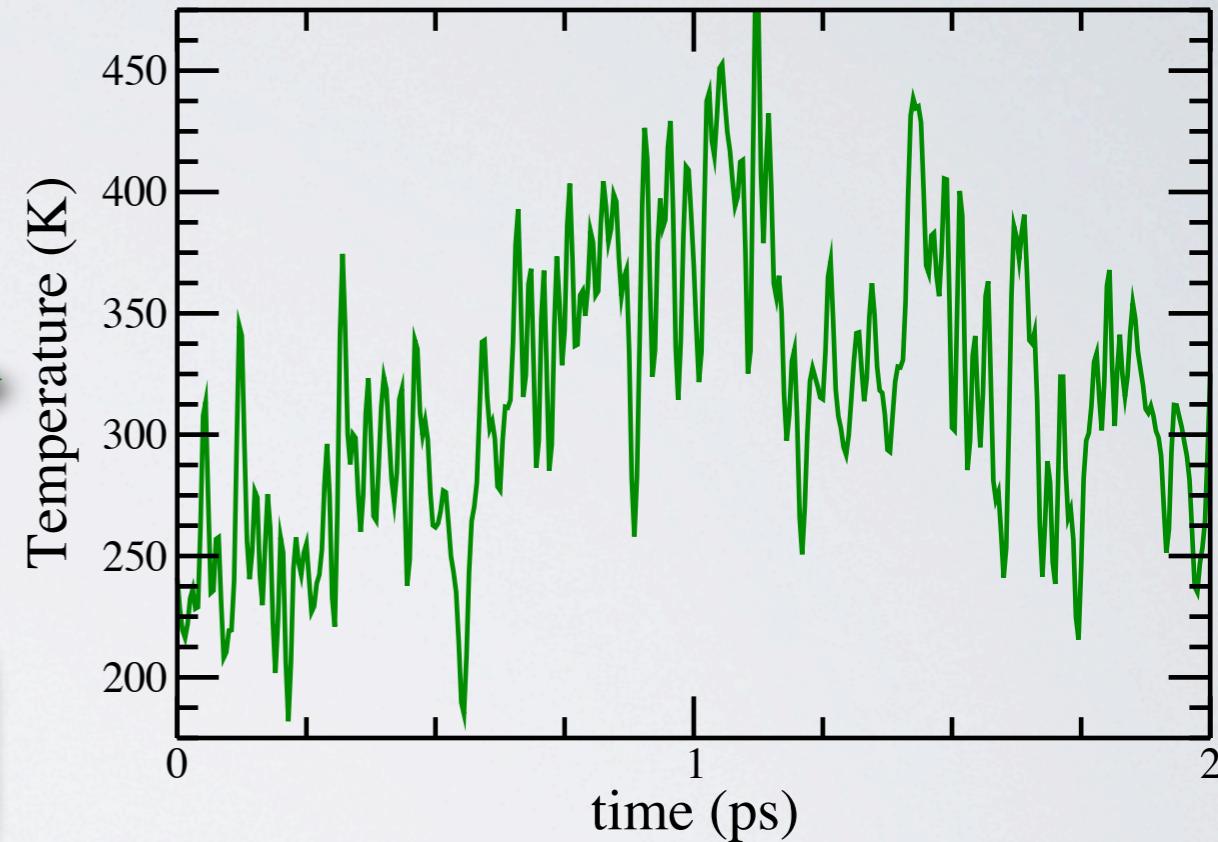
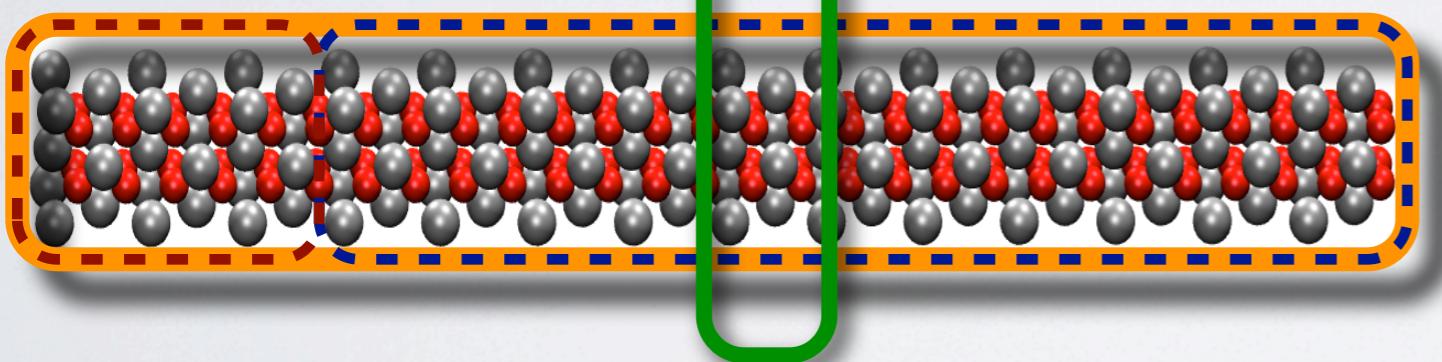


(2) HEAT DIFFUSION

T. M. Gibbons and S. K. Estreicher, *Phys. Rev. Lett.* **102**, 255502 (2009).

Monitor $T(x, t)$

during **thermal equilibration**

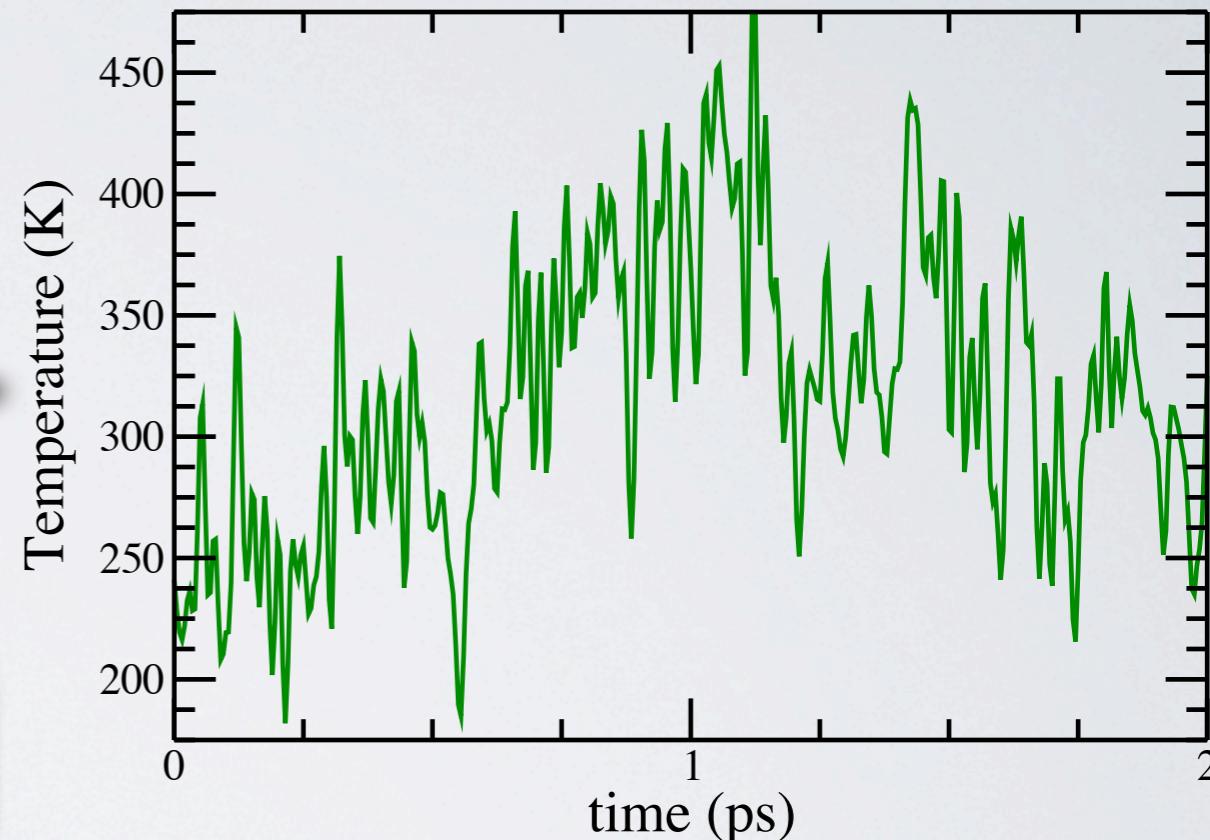
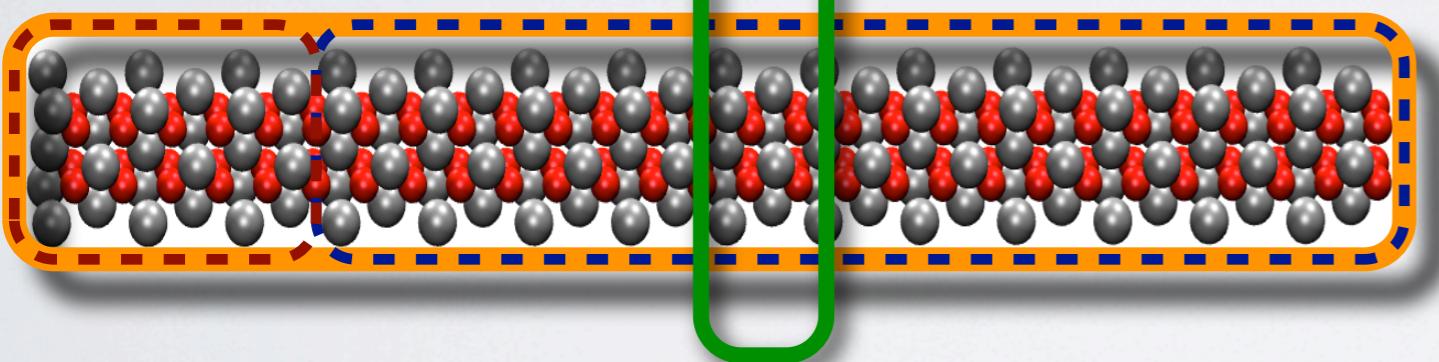


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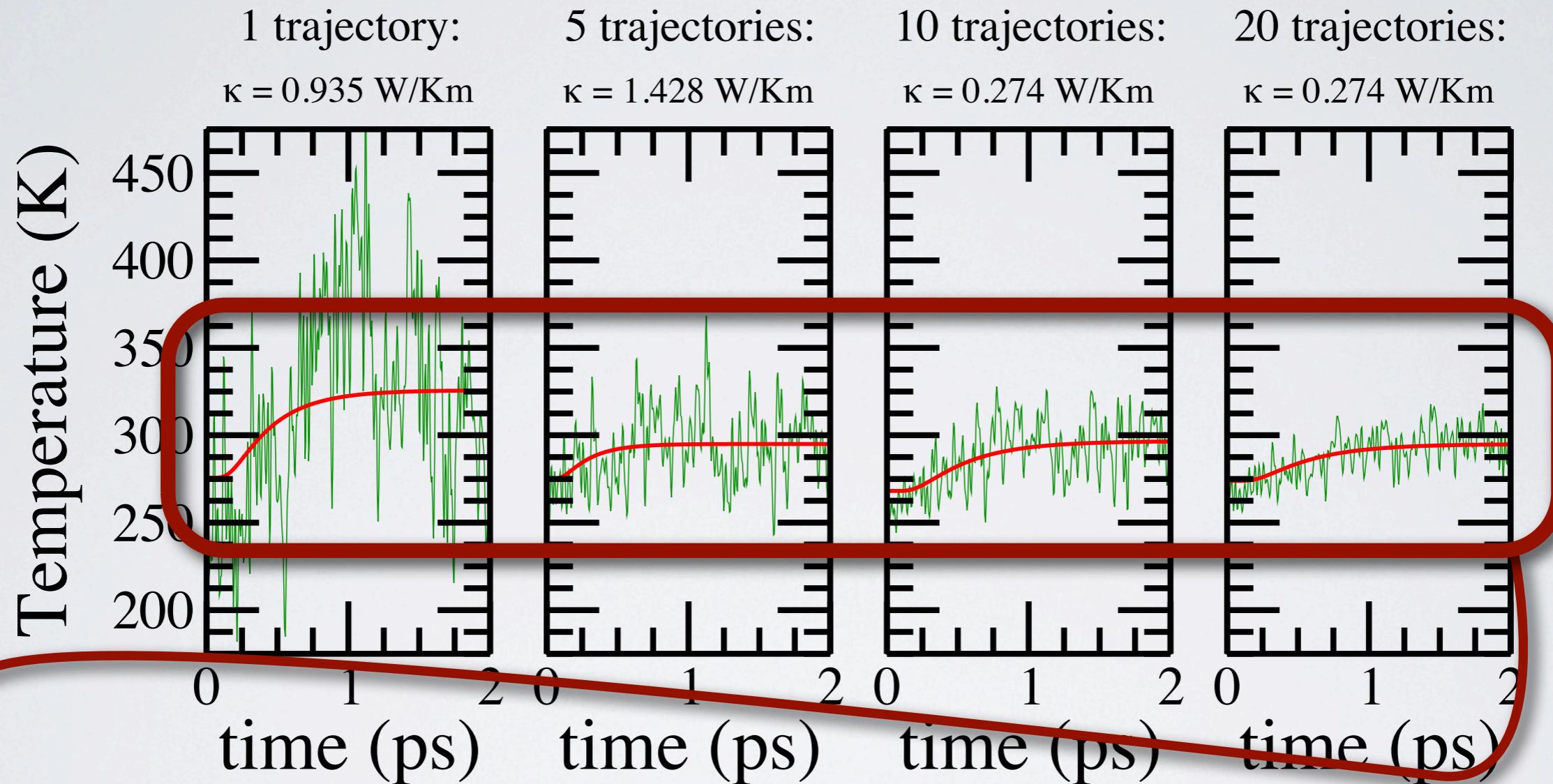
during **thermal equilibration**



Finite supercell size leads to large
temperature fluctuations.

(2) HEAT DIFFUSION

T. M. Gibbons and S. K. Estreicher, *Phys. Rev. Lett.* **102**, 255502 (2009).



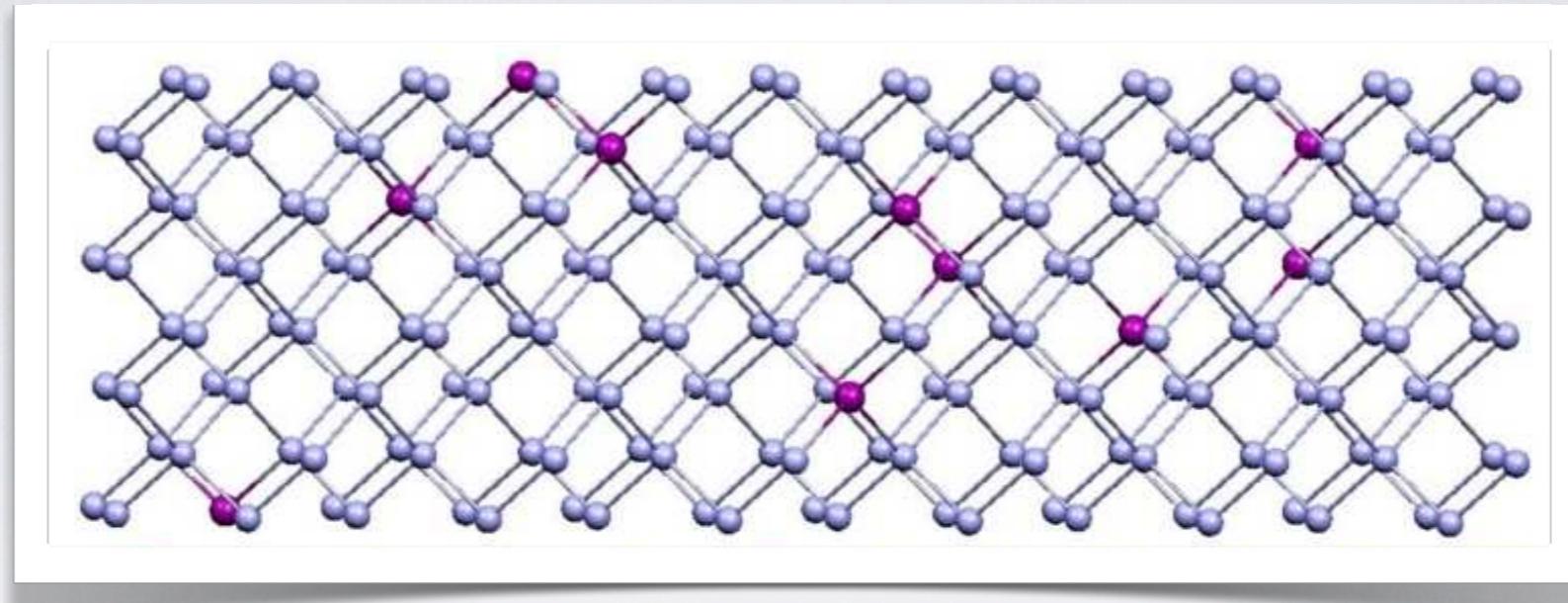
Fit to $T(x, t) = T_{\text{cold}} + (T_{\text{final}} - T_{\text{cold}}) \sum_n (-1)^n \exp\left(-\frac{n^2 \pi^2 \kappa t}{\rho C_V}\right)$

PROS & CONS

- Standard MD method
- No heat source & sink needed
- Minute temperature gradients
- Only applicable at **low** temperatures

(3) APPLICATION TO IMPURITIES IN SI

T. M. Gibbons, By. Kang, S. K. Estreicher, and C. Carbogno, *Phys. Rev. B* (accepted).

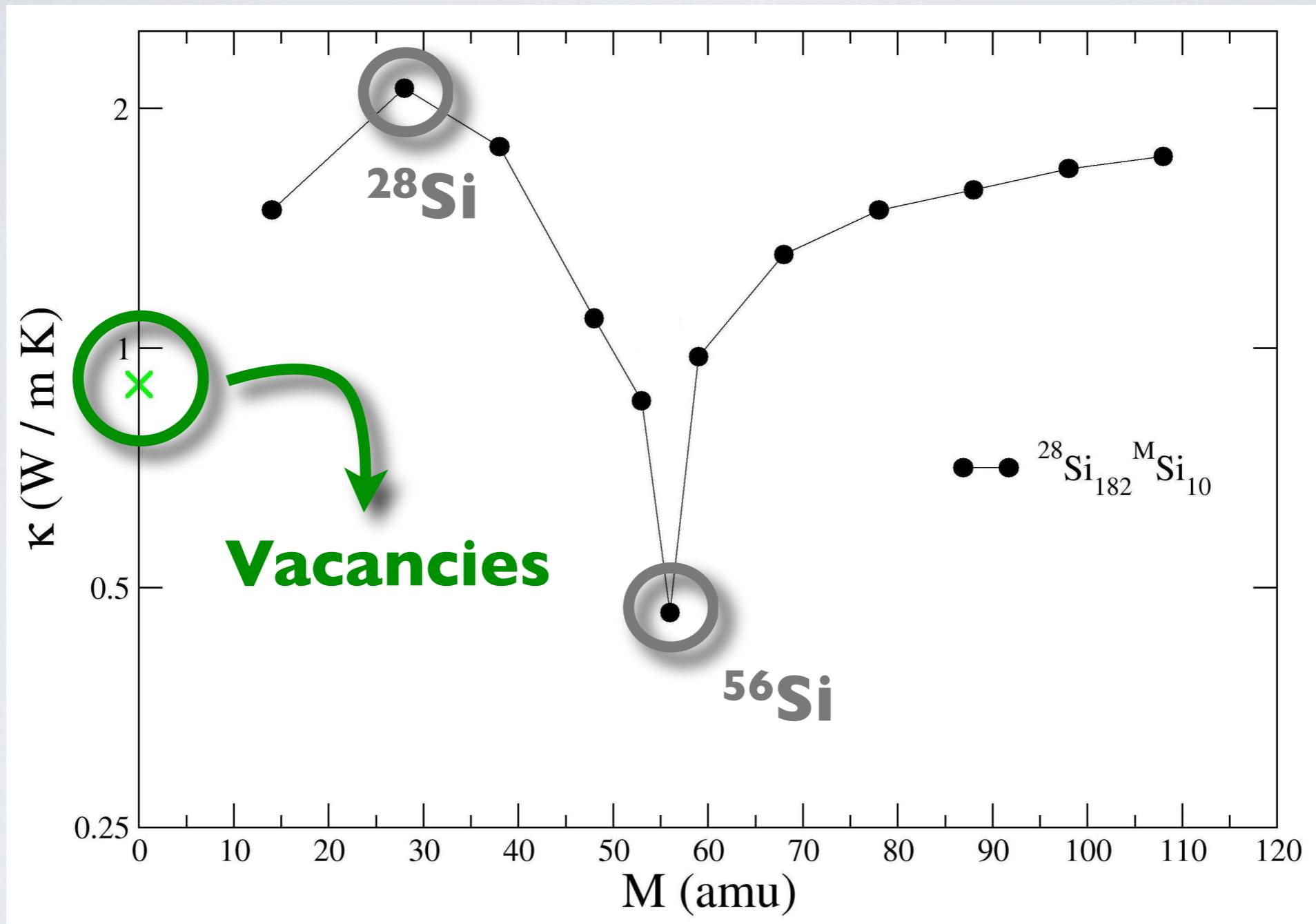


Si₁₉₂ supercell containing **~5.2% impurities**

How do the
properties of the impurities
affect the
thermal conductivity of the system?

(C) APPLICATION TO IMPURITIES IN SI

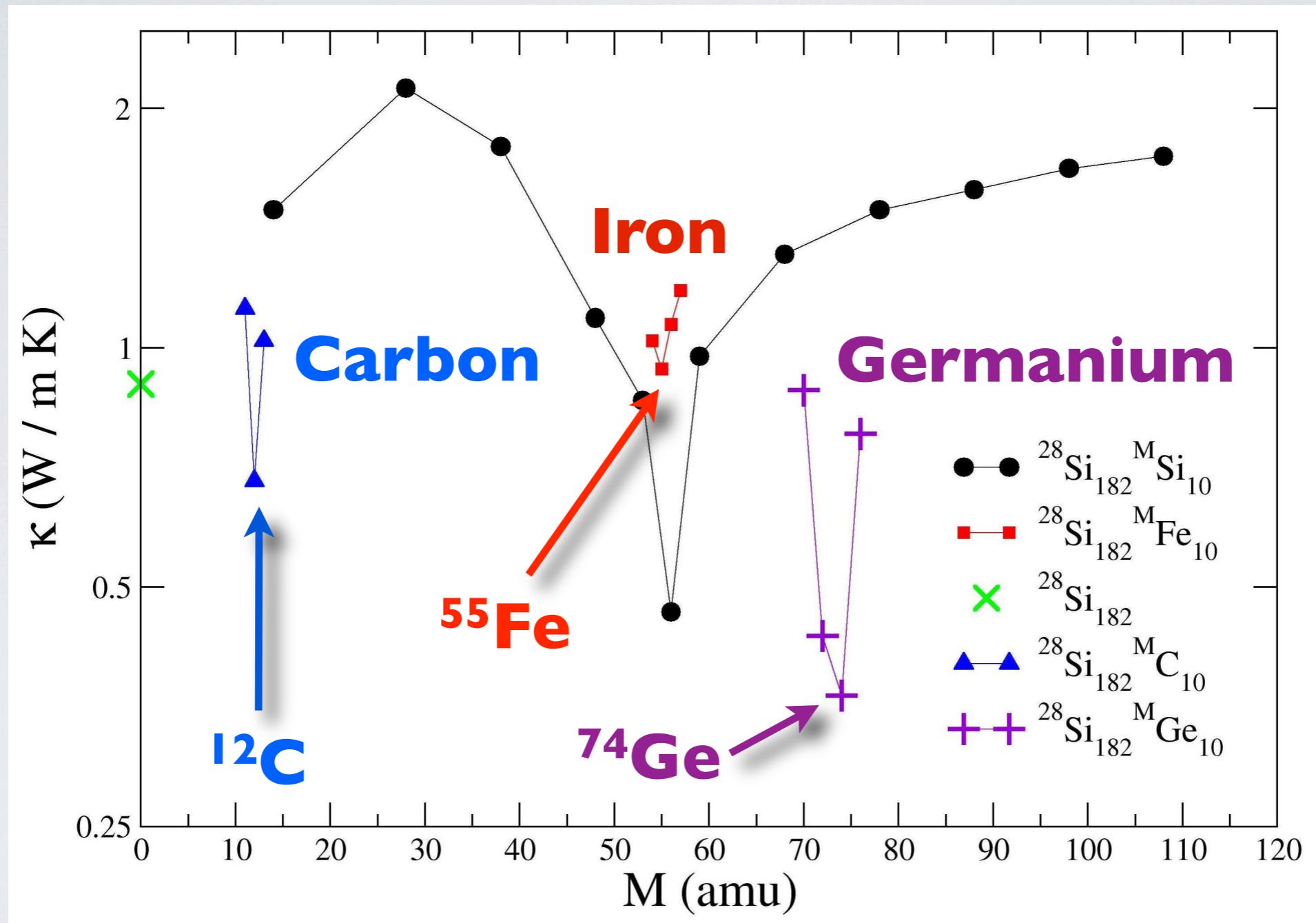
T. M. Gibbons, By. Kang, S. K. Estreicher, and C. Carbogno, *Phys. Rev. B* (accepted).



Thermal conductivity can be controlled via the impurities' mass!

(C) APPLICATION TO IMPURITIES IN SI

T. M. Gibbons, By. Kang, S. K. Estreicher, and C. Carbogno, *Phys. Rev. B* (accepted).



Not all impurities are created equal!

CONCLUSION AND SUMMARY

Molecular Dynamics:

Application in Thermodynamics, Thermostats, and Caveats

Harmonic Approximation:

Basic concepts, Application: Heat Capacities & Lattice Expansion

Thermodynamic Integration:

Complete(!) Anharmonic Free Energy from First Principles

Non-equilibrium dynamics:

Thermal Conductivity from First Principles